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Geodätisches Institut

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GPS/GALILEO-Technologien



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# Preprocessing of high rate GPS data for real-time applications

Diplomarbeit im Studiengang  
**Geodäsie und Geoinformatik**  
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# Table of contents

Zusammenfassung .....	I
Abstract.....	II
1. Introduction.....	1
1.1 Motivation .....	1
1.2 Objective.....	1
1.3 Real-time GPS .....	2
2. Background.....	4
2.1 Cycle Slips.....	4
2.2 Function Model and Definitions.....	4
2.3 Linear combination of data.....	5
2.3.1 Wide-Lane phase combination .....	6
2.3.2 Ionospheric combination .....	7
2.3.3 Ionosphere-free combination.....	8
2.3.4 Differences between Satellites.....	9
2.4 Cycle slip detection in TurboEdit.....	10
2.4.1 Cycle slip detection with the wide-lane combination.....	10
2.4.2 Cycle slip detection in the ionospheric combination.....	11
3. Preprocessing in EPOS .....	13
3.1 Program flowchart .....	13
3.2 Some difference between TurboEdit and TB_Clean .....	15
3.3 Limitation by TB_Clean to deal with real-time data.....	17
4. Compare 1 Hz and 0.033 Hz GPS data .....	18
4.1 Real time 1 Hz GPS streaming data .....	18
4.2 Linear combinations in 1 Hz GPS data.....	18
4.3 1Hz GPS data problem of three IGS stations .....	20
4.4 Single difference between satellites.....	24
5. Detection with single differences .....	28
5.1 Input file and data streams.....	28
5.2 Structure of the software.....	29
5.3 Initialization phase.....	31
5.4 Polynomial Fitting .....	31
5.5 Single difference checking .....	36
5.6 Results .....	37
6. Analysis of single frequency receivers .....	42
6.1 Determination of reference satellite.....	42
6.2 GARMIN (GPS 17-HVS).....	42
6.3 THALES (AC-12) .....	44
6.4 Novatel (smart antenna <sup>TM</sup> ).....	45
7. Summary and conclusion.....	48

# List of figures

Figure 2. 1 Satellite elevation .....	6
Figure 2. 2 Wide-lane ambiguity .....	7
Figure 2. 3 Ionospheric combination .....	8
Figure 2. 4 Ionosphere-free combination.....	9
Figure 2. 5 Cycle slip detection with wide-lane combination .....	11
Figure 3. 1 Program diagram of TB_Clean .....	14
Figure 3. 2 Cycle slip detection with wide-lane in TB_Clean.....	15
Figure 3. 3 Compare all detections in subroutine 'LC_help' .....	17
Figure 4. 1 Ionospheric combination of 1 Hz and 0.033 Hz .....	19
Figure 4. 2 Wide-lane ambiguities of 1 Hz and 0.033 Hz .....	19
Figure 4. 3 Ionosphere-free combinations of 1 Hz and 0.033 Hz.....	20
Figure 4. 4 Wide-lane ambiguities of 1Hz.....	21
Figure 4. 5 Ionospheric combinations of 1 Hz .....	21
Figure 4. 6 Carrier phase $L_1$ and $L_2$ of 1Hz.....	22
Figure 4. 7 Second derivative of $L_1$ and $L_2$ .....	23
Figure 4. 8 Periodogram of $L_1$ and $L_2$ and their difference .....	24
Figure 4. 9 Two satellites elevation.....	25
Figure 4.10 $L_1$ single difference between PRN2 and PRN4.....	25
Figure 4.11 $L_2$ single difference between PRN2 and PRN4.....	26
Figure 4.12 Satellites clock error of PRN2 and PRN4 .....	26
Figure 5. 1 Main program diagram of chn_filter .....	30
Figure 5. 2 Elevation angles of two satellites (top) $L_1$ single difference polynomial fitting (middle) and residual (bottom) .....	33
Figure 5. 3 Outlier detection in polynomial fitting.....	34
Figure 5. 4 outlier detection with 1 order and 2 order polynomial fitting .....	35
Figure 5. 5 2 order polynomial fitting without outliers .....	35
Figure 5. 6 check current epoch.....	36
Figure 5. 7 Check big noises with ionospheric combination.....	37
Figure 5. 8 Variations in combination of 1 Hz GPS data .....	38
Figure 5. 9 RMS after polynomial fitting of 1 Hz GPS data .....	38
Figure 5.10 RMS and elevation angle .....	39
Figure 5.11 RMS after polynomial fitting for all satellites in one day .....	40
Figure 6. 1 Receiver clock errors, ionospheric and troposphere delay.....	43
Figure 6. 2 RMS for GARMIN 1 Hz data checking.....	44
Figure 6. 3 RMS for THALES 1 Hz GPS data checking .....	45
Figure 6. 4 RMS for Novatel 1 Hz GPS data checking .....	46
Figure 6. 5 RMS after polynomial fitting for all satellites .....	47

# List of tables

Table 3.1 Used flags in TB_Clean.....	16
Table 5.1 Input data for preprocessing .....	28
Table 5.2 IGS orbit product in SP3 format.....	29
Table 5.3 Rate of good data after checking (station: PIE1).....	41
Table 6.1 One epoch 1Hz GPS data from GARMIN .....	42
Table 6.2 Rate of good data after checking (THALES) .....	45
Table 6.3 Rate of good data after checking (Novatel).....	46



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# Zusammenfassung

In den letzten Jahren immer mehr geodätische Anwendungen erforderten GPS Daten mit 1 Hz sampling rate, zum Beispiel die Low Earth Orbiter Missionen CHAMP und GRACE sowie GPS Überwachungsnetze. In der gleichen Zeit verbesserten die wachsenden Kapazitäten des Internets den Transport von Daten in Echtzeit und die langfristige Archivierung mit weniger Kosten. Dies erfordert GPS Software mit der Fähigkeit GPS-Daten und-Netze in Echtzeit zu verarbeiten. Weil nur ein Cycle Slip die GPS-Lösung verfälschen kann, ist das Preprocessing von GPS Daten besonders wichtig für die Erhaltung der hohen Genauigkeit.

Zurzeit gibt es eine ganze Reihe von verfügbaren Techniken für diesen Zweck. In dieser Arbeit wird ein neuer Algorithmus vorgestellt, mit dem Cycle Slip von 1 Hz GPS Daten in Echtzeit erkannt werden können. Um die Verarbeitung der 1 Hz GPS Daten in Echtzeit zu gewährleisten, muss der neue Algorithmus schnell und voll-automatisch arbeiten. Desweiteren müssen die Verfügbarkeit, Zuverlässigkeit und Effektivität auch berücksichtigt werden. Nach dem Studium der 1Hz GPS Daten und existierende Algorithmen wird ein neuer Algorithmus in dieser Arbeit entwickelt. Der Algorithmus verwendet Differenzen zwischen Satellitenträgerphasen und ionosphärische (geometriefreie) lineare Kombination. Mit Polynom Fitting werden cycle slips gesucht.

In dieser Arbeit wird die Leistung des Algorithmus zur Bestimmung von Cycle Slips als Funktion von Satellitenelevationswinkeln, anhand Einzelfrequenzempfängern wie Garmin, Thales und Novatek untersucht.

## **Schlüsselwörter:**

- Cycle-Slip-Detektion in Echtzeit
- Real-Time GPS
- Polynom Fitting
- Ionosphärische Linearkombination
- Einzelfrequenzempfängern

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# Abstract

In the past several years, ever more geodetic applications require GPS data with 1-Hz sampling rate, for example, the Low Earth Orbiter missions CHAMP and GRACE GPS monitoring networks. In the same time, the growing capacities of the Internet allowed improvements in the real-time transport of data and long-term archiving became less expensive. This requires GPS software with the capability to deal with real-time GPS data and networks. To obtain high accuracy GPS solution the GPS carrier phase measurements are used. Since even a single cycle slip in GPS carrier phase measurements significantly falsifies GPS solutions, the GPS preprocessing is very important for maintaining the high accuracy.

Currently there are quite a number of available techniques for the pre-processing. But to process high rate GPS data in real time the new algorithm must be able to run in real-time and fully automatically. At the same time the availability, reliability and efficiency must be also taken into account. After studying high rate GPS data and existing algorithm, a new algorithm is developed in this thesis. The algorithm uses single differences between-satellite carrier phases and ionospheric combination (geometry-free linear combination), and applies polynomial fitting to detect cycle slips.

In this thesis the performance of the algorithm will be investigated as a function of satellite elevation angles and the utility of the cycle slip detection for single frequency receivers Garmin, Thales and Novatel is investigated.

**Key words:**

- Cycle-slip detection in real time
- Real time GPS
- Polynomial fitting
- Ionosphere combination
- Single frequency receiver

# 1. Introduction

## 1.1 Motivation

An earthquake is the result of a sudden release of energy in the Earth's crust that creates seismic waves. In 2004 an earthquake in the Indian ocean triggered a series of devastating tsunamis along the coasts of most landmasses bordering the Indian Ocean, killing more than 225,000 people in eleven countries and staggering total of economic damage to US\$ 103 billion [Wikipedia]. With the support of Federal Ministry of Education and Research (Germany) a new project GPS calls GSEIS (Surface Deformations with In Seconds) began in Section 1.1 GeoForschungsZentrum Potsdam (GFZ) on April 1st 2007. The project GSEIS focuses on development and test a system for real-time monitoring of the Earth's surface using GPS/GNSS (Global Navigation Satellite System) technique. Its objective is monitoring effectively earthquakes and volcanic eruptions with the help of satellite remote sensing.

Most of the GPS monitoring systems process data daily or hourly to achieve station coordinates with a precision of a few millimeters. However, in early warning systems the time lag between a natural hazard event and issue of the warning must be as short as possible. In order to quickly estimate the scale of an ongoing rupture GPS data should be streamed via a network, processed in near-real or real-time and integrated with the seismic observations. The importance of the real-time approach should not be underestimated. The real-time strategy is of great significance for warning of volcano eruption, earthquake and tsunami disasters. It has been demonstrated already that dense high-rate GPS networks can measure large dynamic and static surface-ground motions with latencies of only few seconds. Thus, further development of real-time processing strategies and software applications in dense high-rate (1-Hz or higher) GPS networks is urgently needed. [Proposal GSEIS]

Before estimating parameters the received GPS data must be cleared from outliers and cycle slips. An important task of the GPS data pre-processing is the detection and possible correction of cycle slips.

## 1.2 Objective

In this thesis we concentrate on the preprocessing of 1 Hz GPS data for real time applications. Because of the increasing number of GPS stations in GPS networks it is becom-

# Chapter 1 Introduction

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ing difficult to run the preprocessing by one central computer in real time. To make the preprocessing more effective the received GPS data can be checked before it is sent from the GPS observing sites to an analyze central. Currently the most cycle slip detection algorithms are based on double- or triple-difference, and they are not able to realize single GPS station data preprocessing. In Blewitt (1990) the undifferenced cycle slip detection and fixing, called TurboEdit, was introduced. In this algorithm the ionospheric and wide-lane combinations are used for cycle slip detection; but since pseudorange are less precise than carrier phase, it confuses cycle slips with noise in wide-lane combinations. The pseudorange measurements in 1 Hz high rate GPS data have larger noise than that in lower rate GPS data.

In this work the TurboEdit algorithm is introduced, after the analysis of 1 Hz GPS data a new strategy is discussed for single GPS station preprocessing in real time.

## 1.3 Real-time GPS

Today double differencing GPS for accurate positioning is the most common data preprocessing method, because common modes errors such as satellite orbital errors, receiver clock errors, satellite clock errors and atmospheric delay can be reduced for long baselines or even eliminated for short baselines. And after double differencing the double difference ambiguities are constrained to integer values.

Currently carrier phase based GPS kinematic positioning systems are primarily based on double differencing data processing approach which is able to provide centimeter to decimeter accurate positional accuracy in real time. [Gao, Y et al. 2004]

During the last 10 years the availability of precise GPS satellite orbit and clock products of the International GPS Service (IGS) has enabled the development of the new positioning methodology 'precise point positioning' (PPP). In PPP the pseudorange and carrier phase observations from a single GPS receiver are undifferenced. The reference receiver is not necessary any more. To compare with DGPS it is logistically simpler and almost as accurate. [Witchayangkoon, B. 2002]

In 2002 the IGS Real-time Working Group (RTWG) was founded to be aimed to the IGS developing real-time infrastructure and processes. The GFZ Potsdam is one of participating agencies of the RTWG. A prototype system will be developed in GFZ Potsdam for the support of precise real-time positioning. The primary products of such a system will be GPS/GNSS station data and satellite orbits and clocks, made available to the user by Internet and other streaming technologies.

## Chapter 1 Introduction

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At present an increasing number of GPS stations can provide real-time data streams. Therefore appropriate corresponding software packages are needed. A related software package, as a prototype of the future real-time GNSS software at GFZ, has been developed recently for both real-time deformation monitoring and providing service for applications based on PPP under the framework of GSEIS project.

The real-time software packages for GSEIS will be capable of

- Providing satellite orbits and clocks in real-time (network solution).
- Kinematic monitoring of station positions in real-time for networks with good real-time data transmission.
- Performing real-time PPP for stations in high-rate data communications based on orbit and clocks from the network solution (in such cases only orbits and clocks have to be transferred to the station). [Rothacher, M et al. proposal GSEIS 2003]

# 2. Background

Before estimating the parameter (of e.g. station coordinates, receiver clock error etc.) in GPS processing, the observations must be cleaned with the preprocessing. In the preprocessing step poor phase observations are marked. There are quite a number of available techniques for this purpose. In this chapter the definition of cycle slips and several linear combinations will be introduced, which are always used for cycle slips detection. At last part of this chapter the automatic editing algorithm for undifferenced GPS data [Blewitt 1990], called TurboEdit, will be discussed.

## 2.1 Cycle Slips

Carrier phase measurements contain an unknown integer number  $b$  of cycles called ambiguity. This integer number has to be identified in order to exploit the full accuracy potential of GPS carrier phase measurements. During the carrier phase measurements, referenced to the first phase measurement the ambiguity can be continued only when the receiver maintains lock on the received GPS signal. But in some situations, the receiver loses lock of the phase lock loop which might be due to poor signal reception; very fast acceleration changes, or shadowing of satellite signals in their path. Such situations cause discontinuities in the phase measurements, called cycle slips. [Keller, W 2005]

Cycle slips can be any size ranging from one to millions of cycles and can occur concurrently and differently on L1 and L2. After a cycle slip a new ambiguity must be re-estimated. There are lots of methods for cycle slips detection such as triple-difference and TurboEdit [Blewitt 1990]. For cycle slip detection usually the observation combinations and differences are applied. In combinations and differences many errors can be eliminated. Invariably a few cycle slips remain undetected, requiring analyst intervention to fully clean up the data.

## 2.2 Function Model and Definitions

Both the carrier phase and pseudorange observables are used for the PPP method. The following GPS carrier phase and pseudorange equations on two carriers  $L_1$  and  $L_2$  are expressed in units of distance.

$$\Phi_1 = \rho - \frac{If_2^2}{f_1^2 - f_2^2} + \lambda_1 b_1 \quad (2.1a)$$

## Chapter 2 Background

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$$\Phi_2 = \rho - \frac{If_1^2}{f_1^2 - f_2^2} + \lambda_2 b_2 \quad (2.1b)$$

$$P_1 = \rho + \frac{If_1^2}{f_1^2 - f_2^2} \quad (2.1c)$$

$$P_2 = \rho + \frac{If_2^2}{f_1^2 - f_2^2} \quad (2.1d)$$

$\Phi_1$   $\Phi_2$  : measured carrier phase scaled to distance (meters) on carriers  $L_1$  and  $L_2$

$P_1$   $P_2$  : P Code Pseudoranges on carrier  $L_1$  and  $L_2$

$f_1$   $f_2$  : carrier frequencies

$\lambda_1$   $\lambda_2$  : carrier wavelengths

$\rho$  : non-dispersive delay (included geometric delay, troposphere delay, clock signatures and identical effects on carrier  $L_1$  and  $L_2$ )

$I$  : ionospheric delay parameter

$b_1$   $b_2$  : phase ambiguities  $b \in \mathbb{Z}$

GPS satellites transmit two carriers,  $f_1 = 10.23 \text{ MHz} \times 154 = 1575.42 \text{ MHz}$  (wavelength  $\lambda_1 \approx 19.0 \text{ cm}$ ) and  $f_2 = 10.23 \text{ MHz} \times 120 = 1227.6 \text{ MHz}$  (wavelength  $\lambda_2 \approx 24.4 \text{ cm}$ ). On the carriers two codes are modulated, a precision (P) code (on both carriers phase  $L_1$  &  $L_2$ ), and a coarse acquisition code C/A (only on  $L_1$ ).

For ease of discussion, some terms are ignored in the equations, like data noise, multipath, phase center effects, and higher order ionospheric effects.

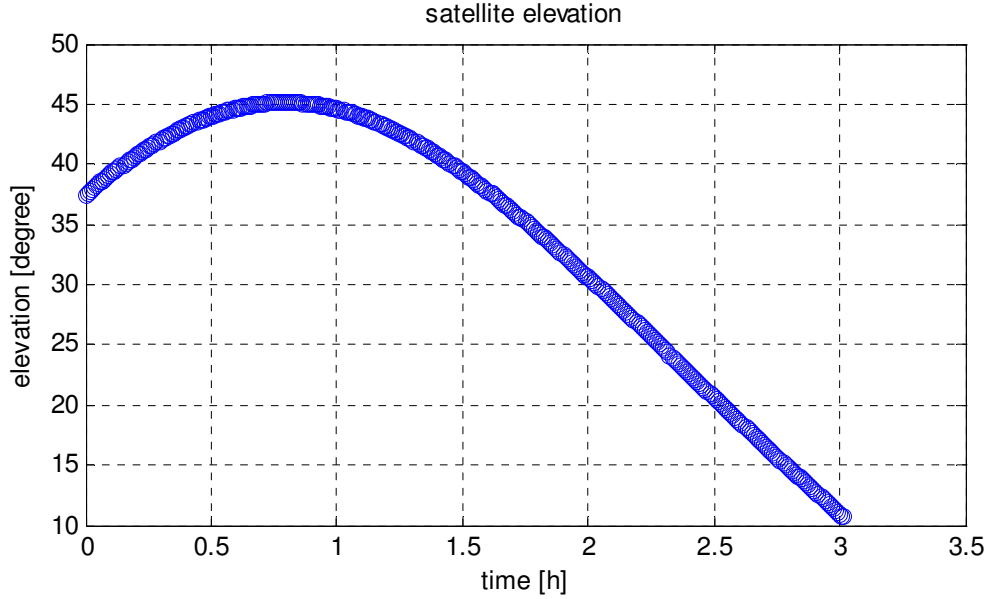
### 2.3 Linear combination of data

Linear combinations and differences of data are very useful, since some errors can be eliminated by combinations and differences. Here three combinations and one difference are introduced.

- Wide-lane Combination
- Ionospheric Combination (geometry-free linear combination)
- Ionosphere-free Combination (ionosphere-free linear combination)
- Single difference between satellites (one receiver and two satellites)

The test data for graphical visualization of the combinations and differences is sampled from the IGS station PIE1 in Pie Town (USA JPL), the sampling rate is 30 seconds. A H-MASER provides the standard frequency. In doy 260 year 2007, the first 3 hours observations from PRN 9 are used.

## Chapter 2 Background



**Figure 2.1 Satellite elevation**

Because cycle slips can occur concurrently and differently on  $L_1$  and  $L_2$ , each signal must be processed independently to check for cycle slips. For these purpose two linear combinations from the equations (2.1a-d) observations can be chosen for cycle slip detection.

### 2.3.1 Wide-Lane phase combination

The wide-lane phase combination and its phase delay and pseudorange can be written:

$$\Phi_w = \frac{f_1\Phi_1 - f_2\Phi_2}{f_1 - f_2} = \rho + \frac{If_1f_2}{f_1^2 - f_2^2} + \lambda_w b_w \quad (2.2a)$$

$$P_w = \frac{f_1P_1 + f_2P_2}{f_1 + f_2} = \rho + \frac{If_1f_2}{f_1^2 - f_2^2} \quad (2.2b)$$

$\Phi_w$  : wide-lane phase expressed as ranges

$P_w$  : wide-lane pseudorange

$b_w$  : wide-lane phase ambiguity  $b_w \equiv b_1 - b_2$   $b_w \in \mathbf{Z}$

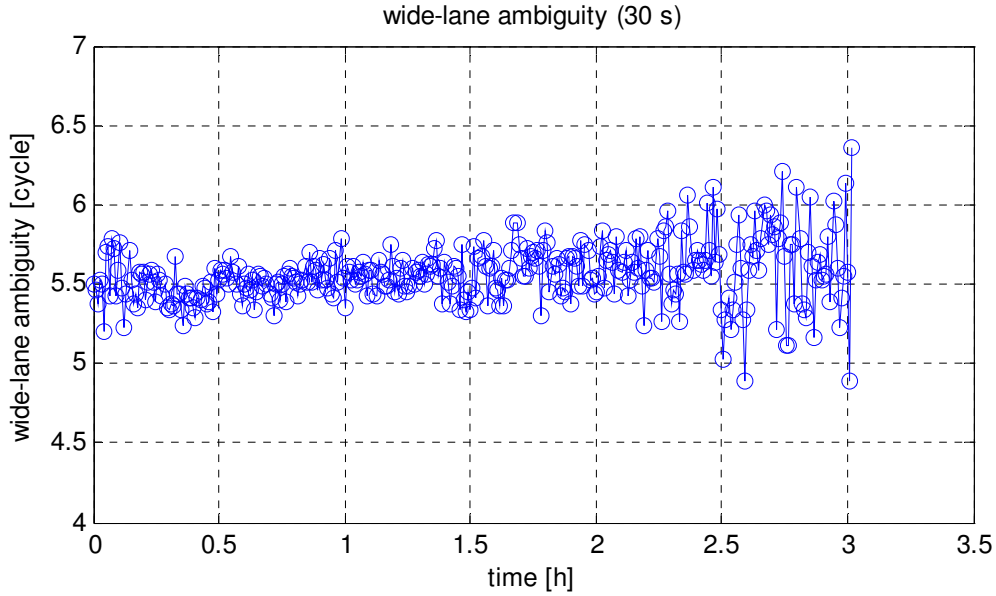
$\lambda_w$  : wide-lane wavelength  $\lambda_w \equiv c / (f_1 - f_2) \approx 86.2$  cm

Because the wide-lane wavelength  $\lambda_w$  is very large, and the noise in the wide-lane pseudorange  $P_w$  is in decimeter level,  $\Phi_w$  can be calibrated by  $P_w$ . From equations (2.2a-b) wide-lane ambiguity  $b_w$  can be presented from pseudorange and carrier phase observation as:

$$b_w = \frac{L_w - P_w}{\lambda_w} \quad (2.2c)$$



## Chapter 2 Background



**Figure 2.2 Wide-lane ambiguity**

Theoretically,  $b_w$  should be constant. Due to multipath and other noise, the wide-lane ambiguity in Figure 2.2 has a variation between 5 and 6. When  $b_w$  in epoch  $i$  has a variation bigger than a given threshold, it is very likely, that a cycle slip has occurred. To find out on which signal the slip is happened, other combination is needed. Here the ionospheric combination (the geometry-free linear combination) is used.

### 2.3.2 Ionospheric combination

The ionospheric delay on GPS signals is frequency-dependent, i.e. the effect on the signal is different on the  $L_1$  and  $L_2$  signals. The ionospheric delay is a function of the latitude of the receiver, the season, the time of day, and the level of solar activity.

The Global Positioning System (GPS) transmits two frequencies, allowing users to correct for the first-order ionospheric signal group delay (or phase advance) that can be as large as 100 meters on  $L_1$  frequency. The second-order ionospheric term, caused by the Faraday rotation effect induced by the Earth magnetic field, is about 1000 times smaller and usually ignored. [Kedar, S et al. 2004]

The ionospheric combination for carrier phase observation is:

$$\Phi_1 \equiv \Phi_1 - \Phi_2 = I + \lambda_1 b_1 - \lambda_2 b_2 = I + \lambda_1 b_\delta - \lambda_1 b_2 \quad (2.3a)$$

$$P_1 \equiv P_2 - P_1 = I \quad (2.3b)$$

$$b_\delta = b_1 - b_2 \quad (2.3c)$$

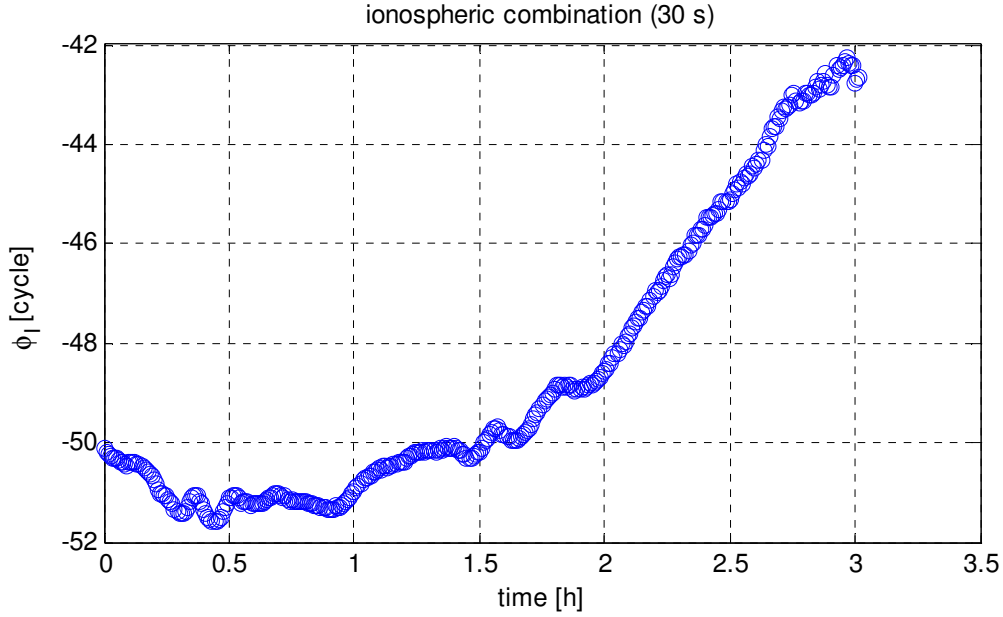
$\Phi_1$  : ionospheric combination carrier phases expressed as ranges

$P_1$  : ionospheric combination Pseudorange

## Chapter 2 Background

$b_\sigma$  : ionospheric combination ambiguity  $b_\sigma \in \mathbb{Z}$

$\lambda_I$  : ionospheric combination wavelength  $\lambda_I \equiv \lambda_2 - \lambda_1 \approx 5.4$  cm



**Figure 2.3 Ionospheric combination**

The pseudorange  $P_I$  is a reasonably smooth function, but because sometimes the multipath effect is bigger than the centimeter level, and the ionospheric combination wavelength is about 5.4 cm, is littler than the wide-lane combination wavelength 86.2 cm,  $\Phi_I$  cannot be successfully calibrated by  $P_I$  like  $\Phi_w$ .

It is possible to construct a polynomial fit  $Q$  to  $P_I$  and subtract it from  $\Phi_I$ , then look for discontinuities as cycle slip. But under high ionospheric activity conditions it is difficult to detect slip with the ionospheric combination, even with high sampling rate.

### 2.3.3 Ionosphere-free combination

The particular ionosphere-free code combination is used on dual-frequency measurements to estimate the first order of the ionospheric delay. For single-frequency receivers it is not possible to account for this ionospheric delay in this way.

The ionosphere-free combinations for codes and phases are defined by:

$$\Phi_c \equiv (f_1^2 \Phi_1 - f_2^2 \Phi_2) / (f_1^2 - f_2^2) = \rho + \lambda_c b_c \quad (2.4a)$$

$$P_c = (f_1^2 P_1 - f_2^2 P_2) / (f_1^2 - f_2^2) \quad (2.4b)$$

$$b_c = b_1 - f_2 b_2 / f_1 \quad (2.4c)$$

$$\lambda_c = f_1 c / (f_1^2 - f_2^2) \approx 48.4 \text{ cm} \quad (2.4d)$$

$\Phi_c$  : the ionosphere-free combination carrier phases expressed as ranges

$P_c$  : the ionosphere-free combination pseudorange

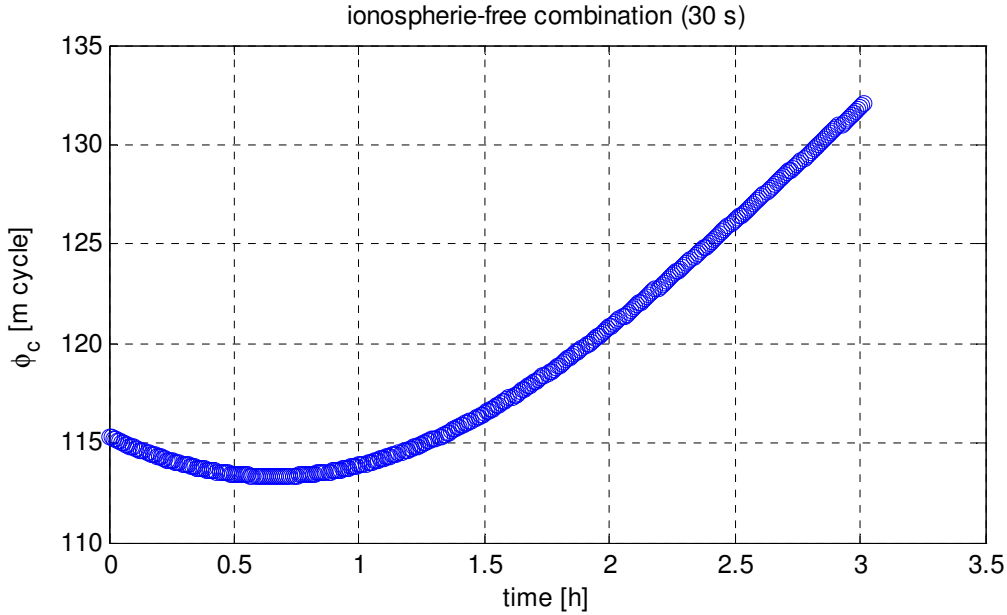
## Chapter 2 Background

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$b_c$  : ionospheric-free combination ambiguity

$\lambda_c$  : ionospheric-free combination wavelength

In the ionospheric-free combination the ambiguity term  $b_c$  is not an integer value any more.



**Figure 2.4 Ionosphere-free combination**

From Figures 2.3 and 2.4 it can be seen, that the values of the ionospheric combination in the 3 hours change less than the values of ionosphere-free combination. But ionosphere-free combination looks like much smoother.

### 2.3.4 Differences between Satellites

For single stations, a useful function for slip detection is the between-satellite single difference (one receiver and two satellites) of carrier phases; it can be presented:

$$d\Phi_1^{m-n} \equiv \Phi_1^m - \Phi_1^n = \rho^{m-n} - \frac{I^{m-n} f_2^2}{f_1^2 - f_2^2} + \lambda_1 b_1^{m-n} \quad (2.5a)$$

$$d\Phi_2^{m-n} \equiv \Phi_2^m - \Phi_2^n = \rho^{m-n} - \frac{I^{m-n} f_1^2}{f_1^2 - f_2^2} + \lambda_2 b_2^{m-n} \quad (2.5b)$$

$d\Phi$  : the phase difference between two satellite m, n

$m, n$  : the satellites indices

The equations (2.5) are free of the receiver clock error. After the discontinuation of SA (selective availability) by The White House at 4.00 UTC on May 2, 2000, the biggest error source in GPS measurement is the receiver clock. The differences between satellites cancel the receiver clock error.

## Chapter 2 Background

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### 2.4 Cycle slip detection in TurboEdit

In 1989 Blewitt has developed an automatic editing algorithm for GPS data for undifferenced cycle slip detection and fixing. In the TurboEdit two linear combinations of the observations are used: the wide-lane combination and the ionospheric combination.

#### 2.4.1 Cycle slip detection with the wide-lane combination

At every observation epoch TurboEdit calculates the wide-lane ambiguity  $b_w$  with equation (2.2c). In addition Blewitt gave a recursive algorithm to update sequentially the wide-lane ambiguity and its variance. For the variance an a priori RMS of half wide-lane cycles is set.

The formulae for the mean value of the wide-lane ambiguity  $\langle b_w \rangle$  and its variance  $\sigma$  is given by:

$$\langle b_w \rangle_i = \langle b_w \rangle_{i-1} + \frac{1}{i}(b_{w,i} - \langle b_w \rangle_{i-1}) \quad (2.6a)$$

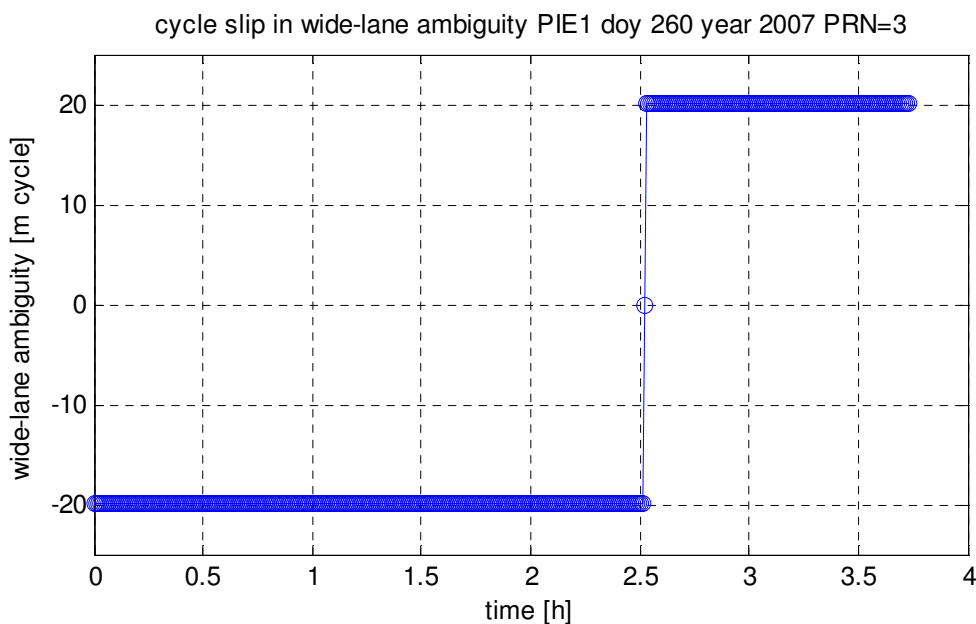
$$\sigma_i^2 = \sigma_{i-1}^2 + \frac{1}{i}[(b_{w,i} - \langle b_w \rangle_{i-1})^2 - \sigma_{i-1}^2] \quad (2.6b)$$

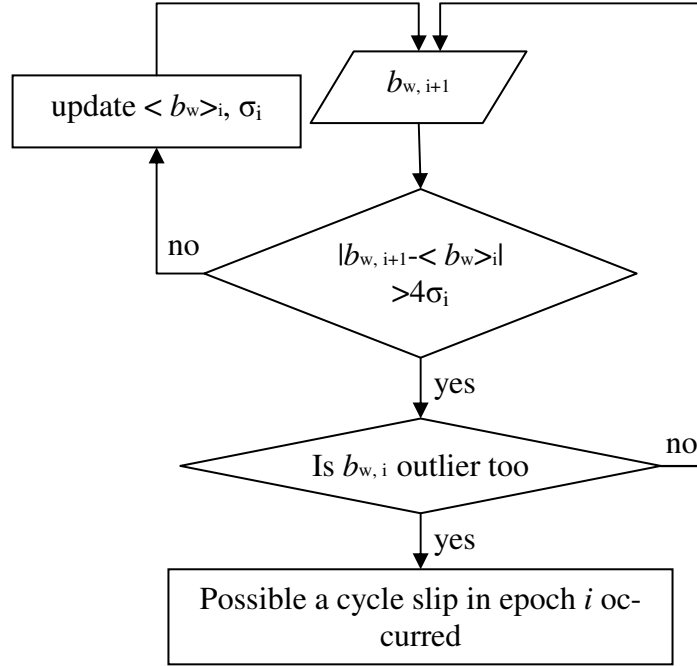
$\langle b_w \rangle$  : mean value of the wide-lane ambiguity  $b_w$

$\sigma$  : variance of the wide-lane ambiguity  $b_w$

$i$  : epoch index in a continued data arc

In equation (2.6a) the mean is calculated exactly; the  $\sigma$  in the equation (2.6b) is an approximation, with a diminishing error term of  $O(1/i^2)$ .





**Figure 2.5 Cycle slip detection with wide-lane combination**

After cycle slip the ambiguity changes to a different integer value, and it begin with a new  $\langle b_w \rangle_i$ . When  $b_w$  of two continued epoch checked out as outliers, it is very possible a cycle slip occurred.

### 2.4.2 Cycle slip detection in the ionospheric combination

In the extremely unlikely case that the cycle slip in  $L_1$  equals the one in  $L_2$ , then the wide-lane slip detection can not detect the cycle slip. In this case the ionospheric combination is used for cycle slip detection.

From equation (2.3b) it can be seen that the ionospheric pseudorange combination  $P_1$  is a reasonably smooth function, but because the multipath in pseudorange measurement is very often bigger than 10 cm, and  $\Phi_1$ 's wavelength is about 5.4cm.  $\Phi_1$  can not be calibrated by  $P_1$ . To calibrate  $\Phi_1$  a polynomial  $Q$  is fitted to  $P_1$  and then subtracted from  $\Phi_1$ . The residuals are searched for discontinuities as outliers or cycle slips.

To choose the degree  $m$  of polynomial fit Blewitt suggested an empirical formula:

$$m = \min[(N/100 + 1), 6] \quad (2.7a)$$

$N$  : number of observations in the data

For detection of discontinuity in  $(L_1 - Q)$  the algorithm is:

$$(L_{1,i} - Q_i) - (L_{1,i-1} - Q_{i-1}) > k \text{ cycle} \quad (2.7b)$$

$$(L_{1,i+1} - Q_{i+1}) - (L_{1,i} - Q_i) < 1 \text{ cycle} \quad (2.7c)$$

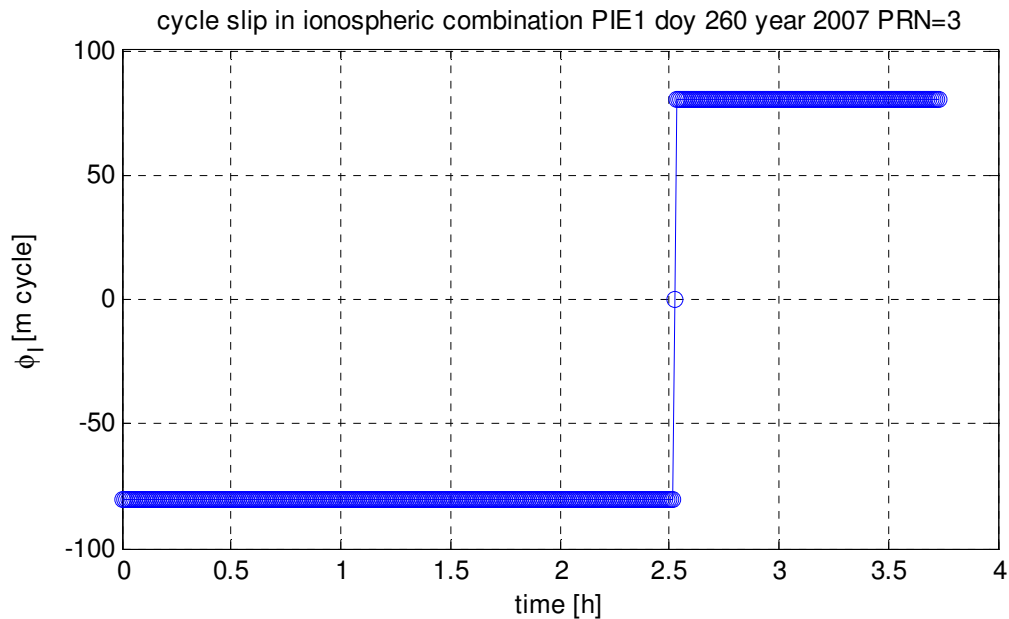
$i$  : epoch index

## Chapter 2 Background

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$k$  : threshold, default value is 6

When the above equations (2.7b) and (2.7c) are satisfied, then the epoch  $i$  is the first good data epoch point after the occurrence of a cycle slip.



**Figure 2.6 Cycle slip detection with ionospheric combination**

Because of ionospheric activity for receivers at high latitudes and most cycle slip bigger than 6, Blewitt suggests: the default value of  $k$  is set to 6 ionospheric cycles ( $6 \times 5.4 \text{ cm} = 32.4 \text{ cm}$ ), but the threshold can be set to a more appropriate value depending on the ionospheric conditions.

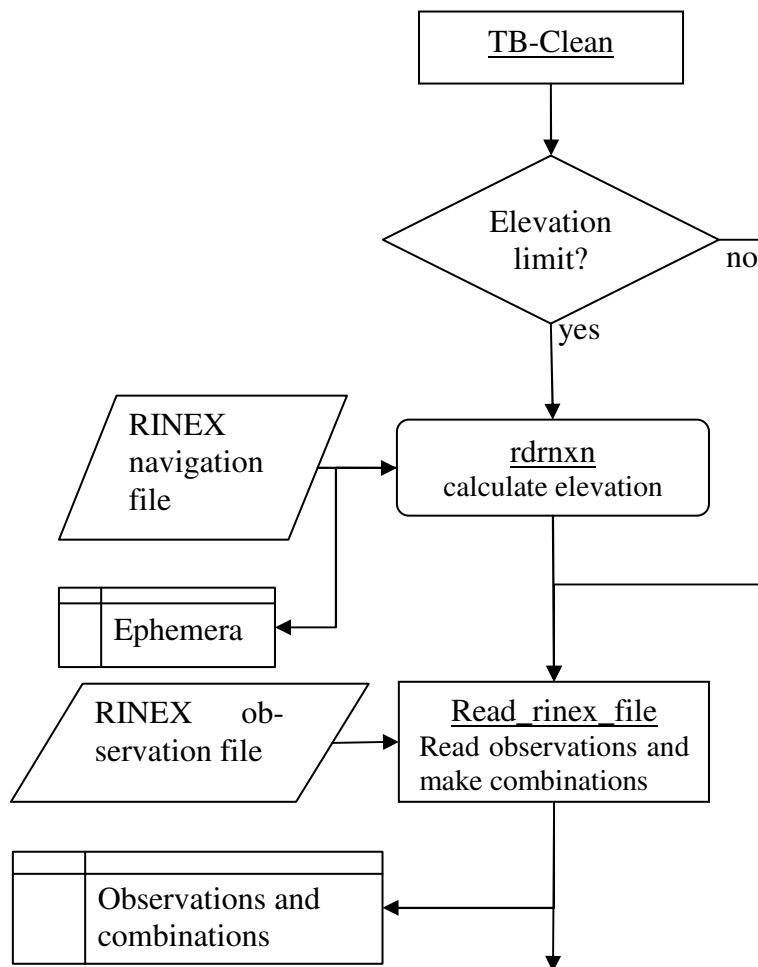
However, there is a hidden problem in the ionospheric carrier phase combination for certain slip combinations. For example, a 24-cycles slip in  $L_1$  and a 19-cycles slip in  $L_2$  cause no jump in the ionospheric phase combination.

### 3. Preprocessing in EPOS

In the latest version of the EPOS (Earth Parameter & Orbit System) [Ge,M. et al. 2005] the preprocessing module “TB\_Clean” is based on the algorithm TurboEdit. The implementation of Blewitt’s algorithm TurboEdit in TB\_Clean is introduced and some improvements are discussed.

#### 3.1 Program flowchart

The TB-Clean is written in Fortran 90. The input files are observation files and navigation files in RINEX format. In the process of TB-Clean all measurements are ticked. To simplify the later parameters estimation the continued epochs longer than 5 minutes are used.



## Chapter 3 Preprocessing in EPOS

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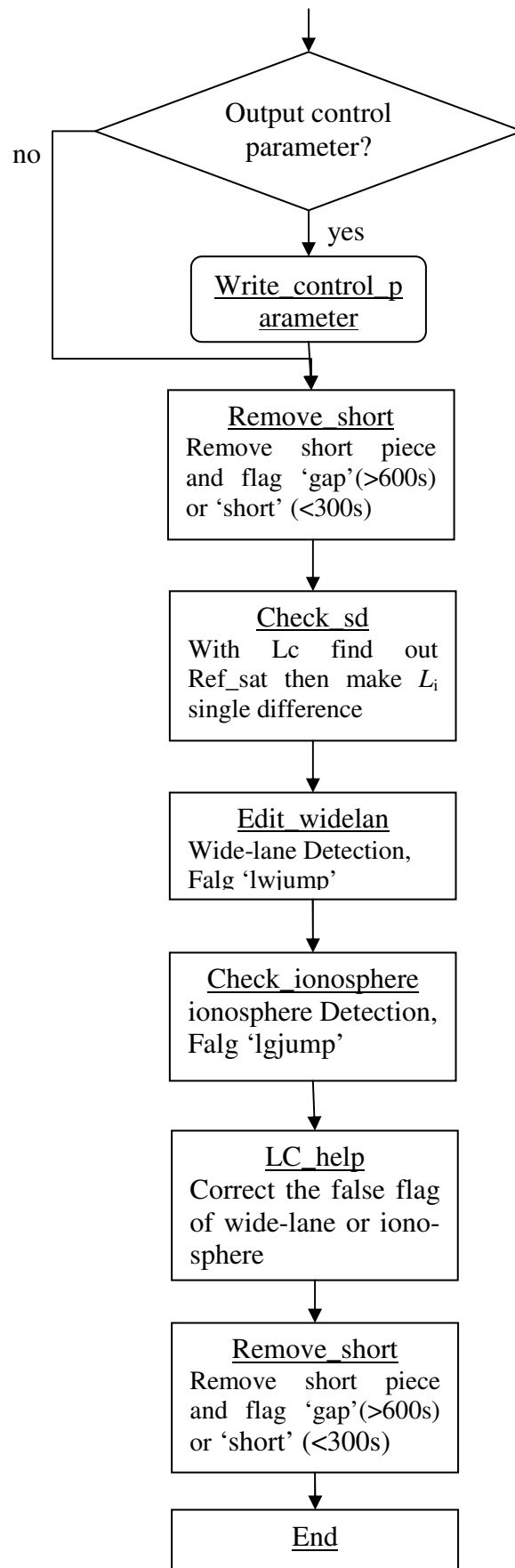


Figure 3.1 Program diagram of TB\_Clean



## Chapter 3 Preprocessing in EPOS

### 3.2 Some difference between TurboEdit and TB\_Clean

The thresholds in the program TB\_Clean are different from those given by Blewitt. As Blewitt published TurboEdit in 1990, the SA (Selective Availability) technique was still active, the SA can introduce intentional, slowly changing random errors of up to a hundred meters into the publicly available navigation signals to confound. After the SA was terminated on May 1, 2000, the single difference between satellites is possible to be used to detect cycle slips and outliers. In TB\_Clean there single difference between satellites ionospheric-free combination is used to detect cycle slips and it is very effective to deal with high ionosphere activity.

a) Threshold for wide-lane phase cycle slips detection.

In the cycle slip detection in the wide-lane combination of Blewitt if  $|b_{w, i+1} - b_w| > 4\sigma_i$ , then the epoch  $i+1$  observation is considered as a cycle slips. From the formulae for the wide-lane ambiguity  $b_w$  we know, that the ambiguity  $b_w$  should be a constant in a arc, when there is no cycle slip. If a cycle slips happens in  $L_1$  or  $L_2$ , theoretically it should appear a same cycle slips by wide-lane ambiguity. In the subroutine TB\_Clean the limit value is set as 1 cycle, one Wide-Lane wave length is about 86.2 cm. And a more condition is given: the jump between wide-lane ambiguities in two continued epochs should be smaller than half cycle.  $|b_{w, i+1} - b_w| < 0.5$  cycle, about 43.1 cm. This condition requires the observations more smooth.

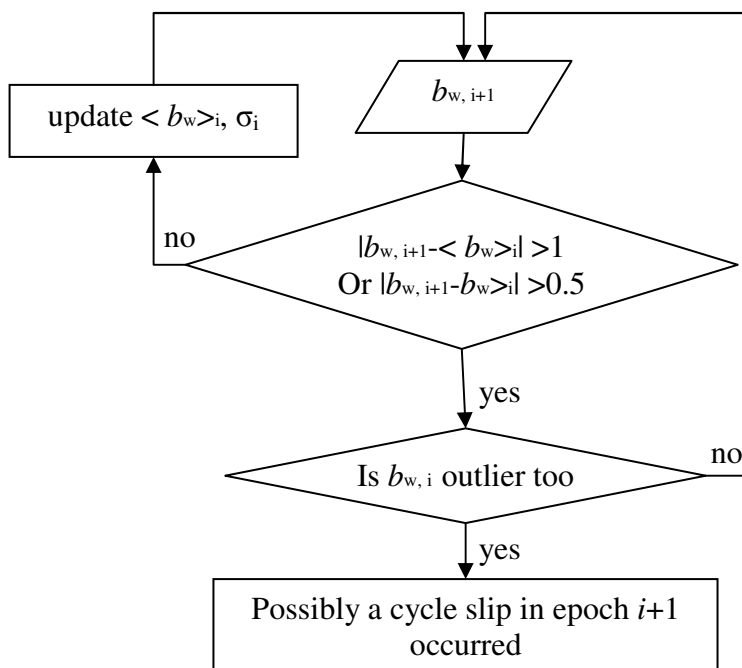


Figure 3.2 Cycle slip detection with wide-lane in TB\_Clean

## Chapter 3 Preprocessing in EPOS

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### b) Ionospheric phase combination cycle slip detection

For the degree  $m$  of polynomial fit  $Q$  of the ionospheric combination is declared with the empirical equation (2.7a) by Blewitt.

But usually a quadratic polynomial fit is enough to deal with normal ionospheric activity; therefore in the subroutine `TB_Clean` the polynomial fit has no more than 2 degrees.

### c) With single difference between in ionosphere-free phase combination

Blewitt also remarks, that under conditions of high ionospheric activity, the ionosphere-free combinations are a good choice for cycle slip detection. But unfortunately the ionosphere-free combinations is very sensitive to clock variations, most clocks of normal GPS receiver have this shortcoming. Then he suggested that it is also effective to difference the ionosphere-free combinations data between satellites.

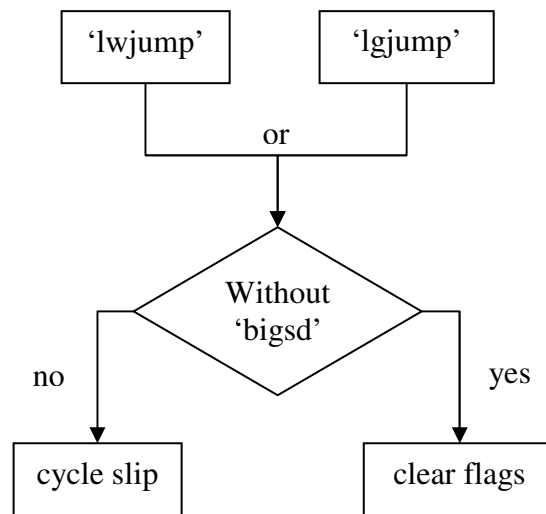
Because the velocity of GPS satellite is very large, the ionosphere-free phase combinations increases or decreases also very quickly, for polynomial fitting it is not easy to find out small cycle slip. To avoid this situation the ionosphere-free phase combination is subtracted from the distance, which can be calculated from GPS broadcast ephemeris and approximate receiver position.

With the ionosphere-free combinations in continued 20 epochs a reference satellite is found out, which has most availed data in the 20 epochs and has a small RMS by a 2 degrees polynomial fit. Then after subtraction of satellite-receiver distance the ionosphere-free combinations of the other satellites are differenced with the reference satellite in these 20 epochs. After the fitting the big discontinuities are flagged as 'bigsd' (big jump in single difference).

Subroutine name:	Used combination	Flag for possible cycle slip
<code>edit_widelane</code>	Wide-Lane Combination	'lwjump'
<code>check_ionosphere</code>	Ionospheric Combination	'lgjump'
<code>check_sd</code>	Ionosphere-free Combination	'bigsd'

**Table 3. 1 Used flags in `TB_Clean`**

d) After the three combinations detection the subroutine '`LC_help`' in `TB_Clean` compare with all flags, and make decision whether the tested epoch includes cycle slips. If the epoch has only one flag of 'lwjump' or 'lgjump', but without flag 'bigsd', then all flags of the epoch are cleared. This means, the epoch is accepted as a normal epoch without cycle slips.



**Figure 3.3** Compare all detections in subroutine `'LC_help'`

### 3.3 Limitation in `TB_Clean` to deal with real-time GPS data

In the subroutine `TB_Clean` to detect cycle slip the polynomial fitting is applied. And the observations not only before and also after the test epoch are used to make the polynomial fitting. For real-time data, it is impossible to use the data after current epoch. Therefore the `TB_Clean` must be adapted to deal with real-time data, or a new method should be developed.

# 4. Compare 1 Hz and 0.033 Hz GPS data

Although many GPS receivers can operate at 1 to 10 Hz, most GPS measurements are sampled at 0.033 Hz. In the past several years, a growing number of continuous operating GPS receivers have increased their recording rates to 1 Hz. The high rate GPS data is beginning to be used for a variety of geophysical monitoring purposes, including seismology, missions in low Earth orbit. All above applications required 1 Hz GPS data processing in real time.

As improvement of the estimates of high-rate GPS position get more precise, the value of these high rate GPS monitoring systems become more important. In this chapter the differences of combinations for 1 Hz GPS data and 0.033 Hz GPS data are discussed, and a new data analysis's is proposed to detect cycle slip.

## 4.1 Real time 1 Hz GPS streaming data

Ntrip (Networked Transport of RTCM via Internet Protocol) is a procedure developed by Bundesamt für Kartographie und Geodäsie (BKG, Frankfurt am Main) and University of Dortmund to provide GNSS correction data as streaming data via the Internet. NTRIP is base on other streaming data procedures, for e. g. Internet Radio. [VermKV (Vermessungs- und Katasterverwaltung Rheinland-Pfalz)]

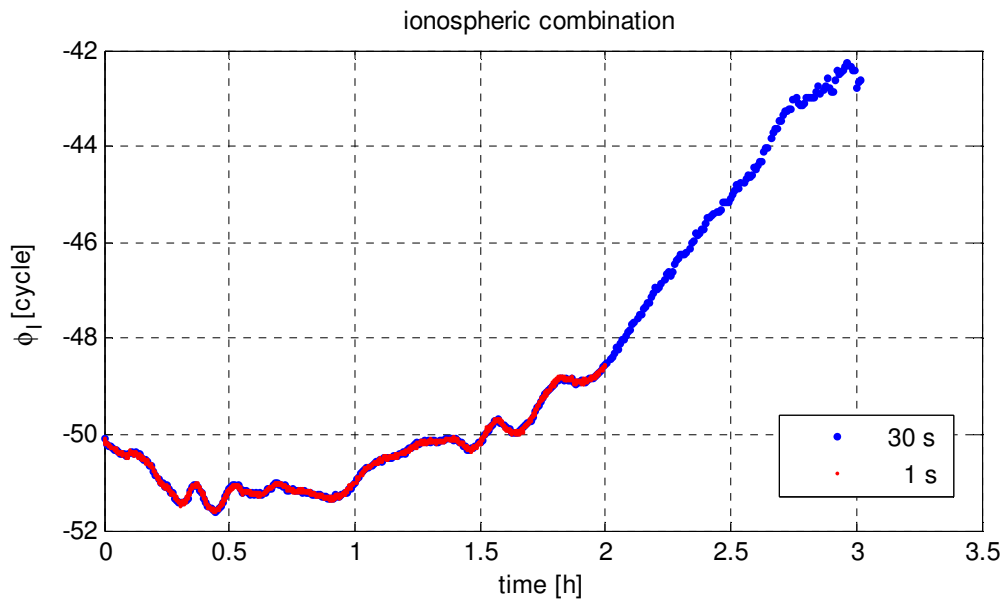
Using the NTRIP data streaming technology, raw data of globally distributed GNSS stations is retrieved from the BKG Ntrip Client by BKG, and are used for the estimation of both satellite clocks for PPP and station positions in network mode. The results are validated with IGS final products and with known station positions.

The project GSEIS gets real time 1 Hz GPS streaming data with the NTRIP data streaming technology. And our preprocessing should to process the streaming data.

## 4.2 Linear combinations in 1 Hz GPS data

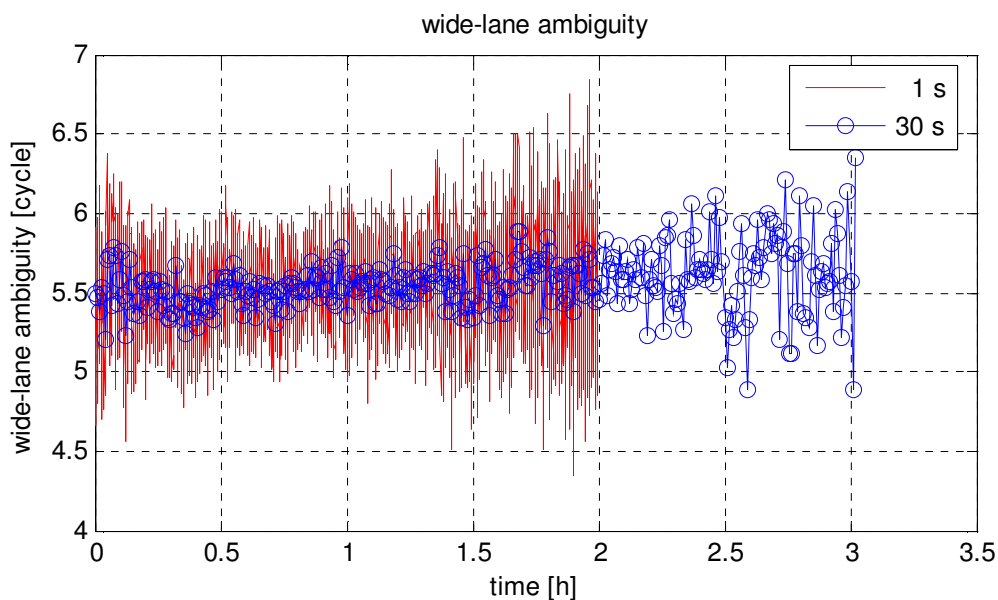
Here the linear combinations in 1 Hz GPS data are compared with the linear combination in 0.033 Hz GPS data.

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



**Figure 4.1 Ionospheric combinations of 1 Hz and 0.033 Hz**

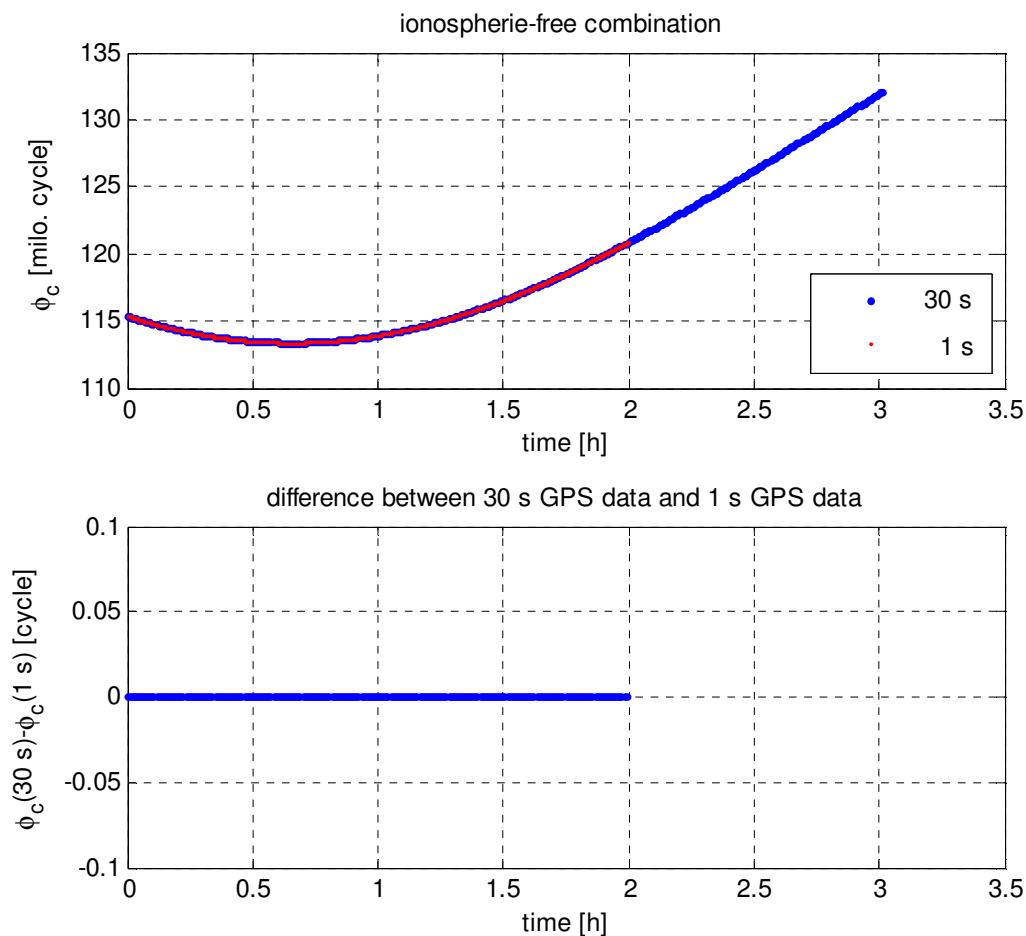
The Figure 4.1 shows us, that there is no difference in ionospheric combination between 1 Hz GPS data and 0.033 Hz. The ionospheric combination can be used to detect cycle slip. But some time because of high ionosphere activity the line is not smooth; in this case it can be difficult for polynomial fitting with low degree.



**Figure 4.2 Wide-lane ambiguities of 1 Hz and 0.033 Hz**

From the identification of ionospheric combination's and the wide lane ambiguity's definition it can be seen, that in Figure 4.2 the bigger variations of 1 Hz wide lane ambiguity should come from the noise of pseudorange measurement. And it is too big to be used for cycle slip detection.

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



**Figure 4. 3 Ionosphere-free combinations of 1 Hz and 0.033 Hz**

The Figure 4.3 confirms again, that the 1 Hz GPS data sampling has same carrier phase measurements as 0.033 Hz GPS data.

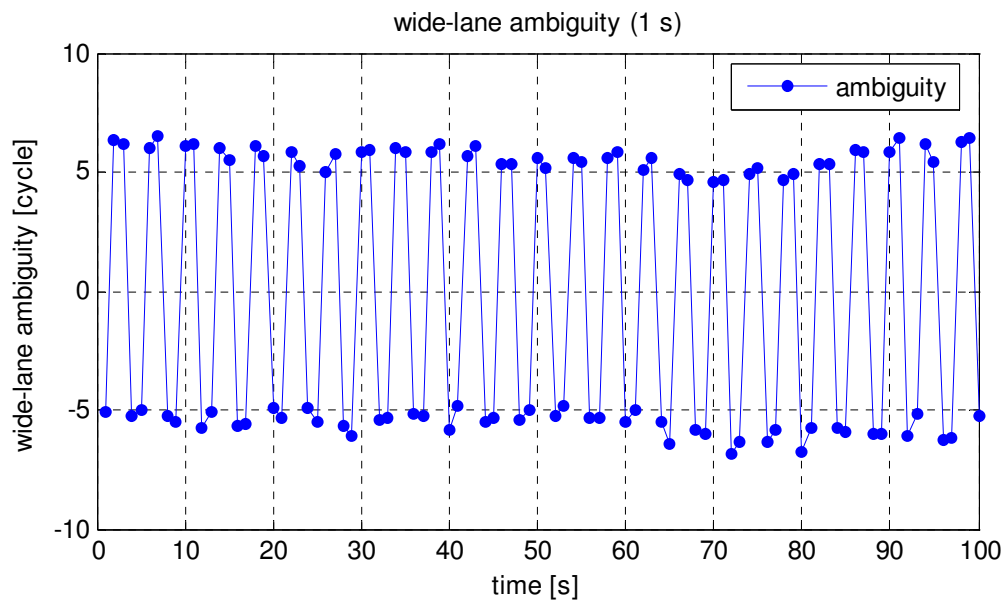
As explained in chapter 3, because the velocity of GPS satellite is very large, for effective polynomial fitting of the ionosphere-free phase combination the distance between satellite and receiver should be subtracted, which can be calculated from GPS broadcast ephemeris and approximate receiver position. And the distance is calculated iteratively. For real-time PPP preprocessing purpose algorithm with iterations should be avoided.

### 4.3 1Hz GPS data problem of three IGS stations

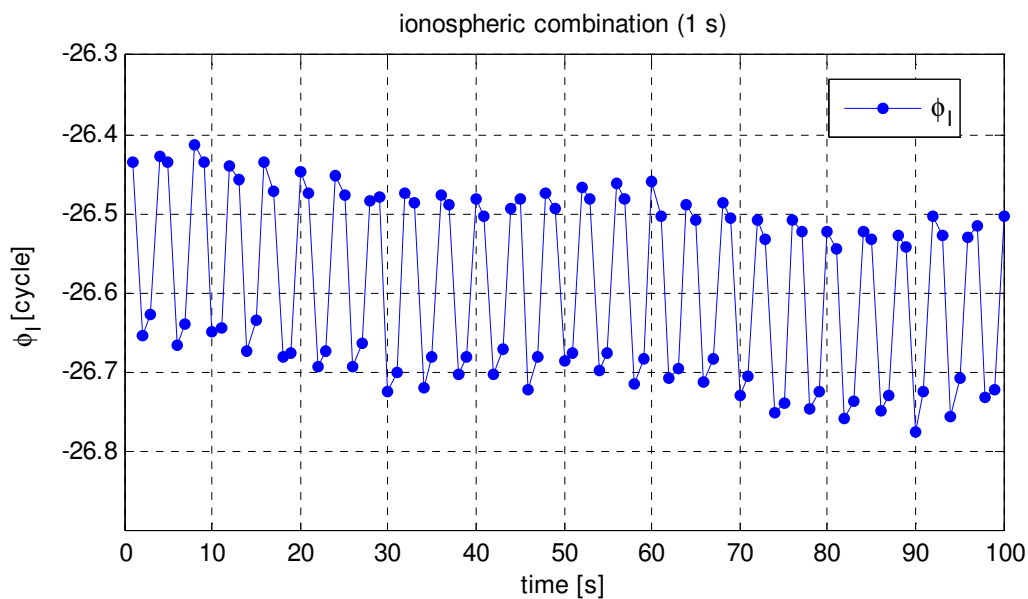
Now about 110 sites of the IGS network can sample and send 1 Hz GPS data. 6 days' GPS data are downloaded. During the test it could be observed, that the IGS stations, POTM in Potsdam, WUH2 in Wuhan and OBE2 in Oberpfaffenhofen, have problems by sampling 1 Hz high-rate GPS data. As shown in the Figure 4.4 to 4.5.

The combinations oscillate at a frequency of about 0.25 Hz.

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



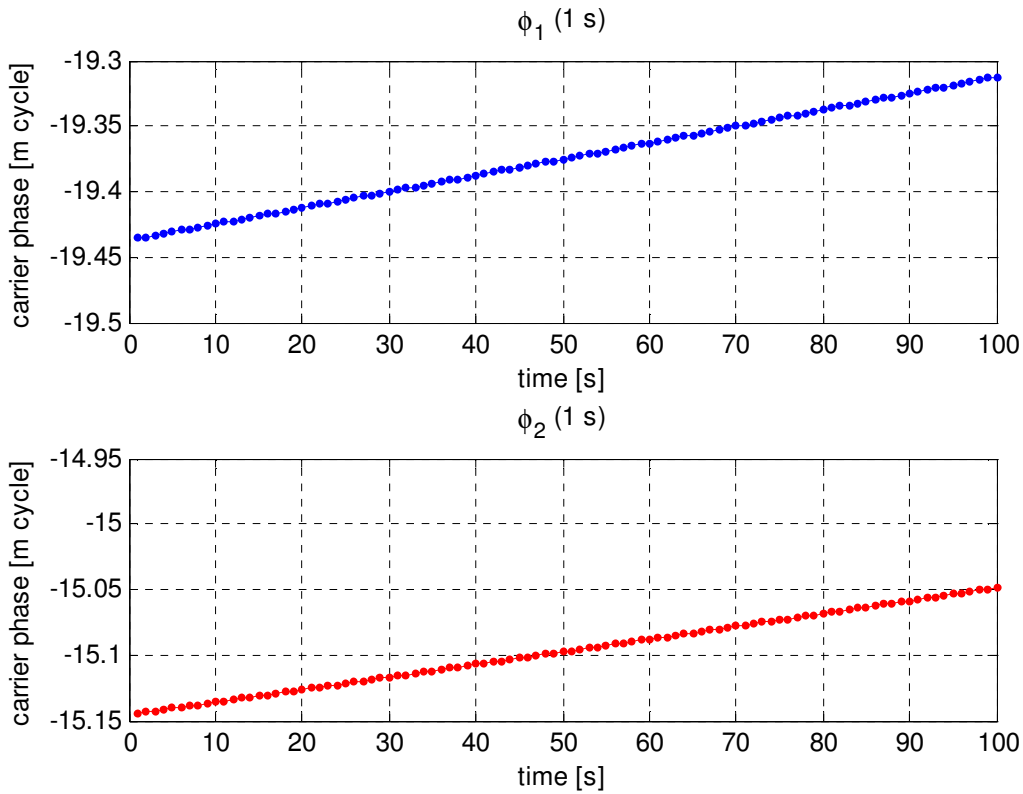
**Figure 4. 4 Wide-lane ambiguities of 1Hz (data from WUH2, doy 260 year 2007 PRN3)**



**Figure 4. 5 Ionospheric combinations of 1 Hz (data from WUH2, doy 260 year 2007 PRN3)**

From the equation (2.3) we know that the ionospheric combination is linear combination from  $L_1$  and  $L_2$ , i.e. one or both of received signals  $L_1$  and  $L_2$  were not correctly generated by the GPS receiver.

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



**Figure 4. 6 Carrier phase  $\Phi_1$  and  $\Phi_2$  of 1Hz (data from WUH2, doy 260 year 2007 PRN3)**

Because the GPS satellites have very high velocity, the carrier phase measurements increase rapidly. In the ionospheric- and wide lane combinations the geometric parts are cancelled, but there are also disadvantages, that the noise of the both the carrier phase measurements are enlarged.

Alone from the above four figures it is not clearly, which carrier phase measurements of WUH2 station causes the oscillation in the ionospheric combination and in wide lane combination.

The relative motion between GPS satellites and site receiver can be approximated by a constant acceleration motion.

Because of the high increase in the carrier phase we can not check the spectrum of the carrier phase  $\Phi_1$  and  $\Phi_2$  directly. Instead of the carrier phase  $\Phi_1$  and  $\Phi_2$  their second derivative are used to check the spectrum.

$$x_p(t) = \frac{A_0}{2} + \sum (A_k \cos k\omega_0 t + B_k \sin k\omega_0 t); k \leq K \quad (3.1a)$$

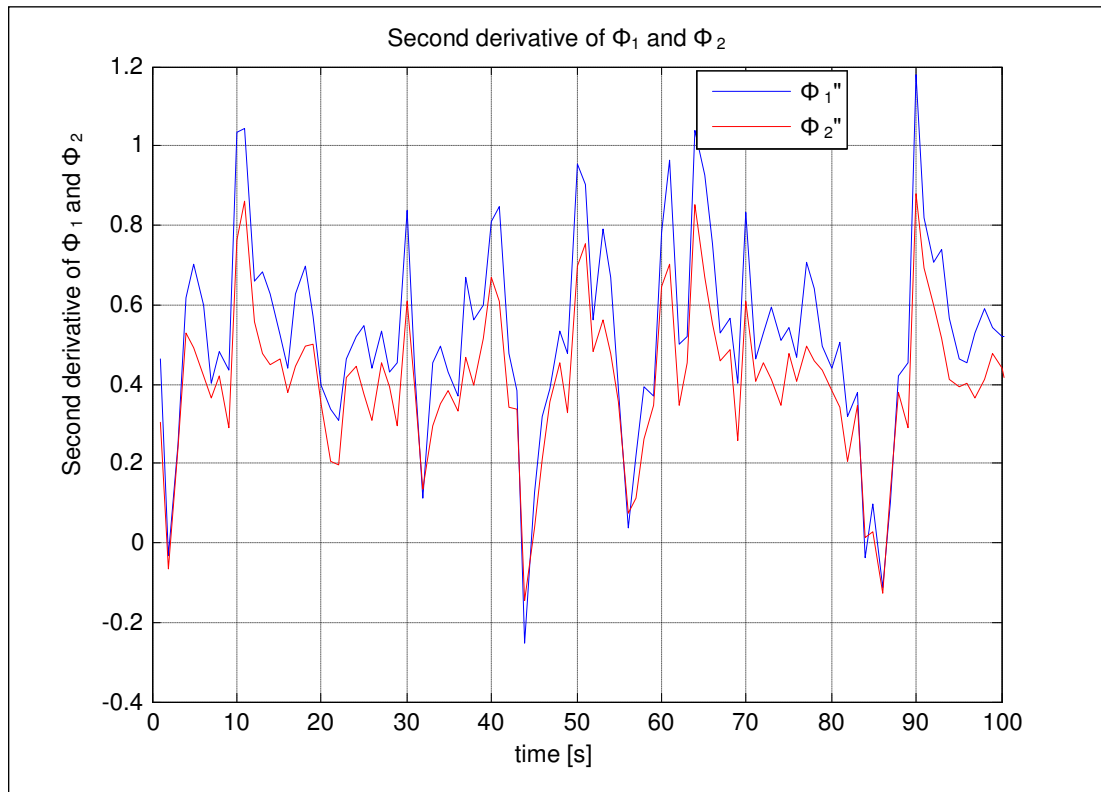
$$x_p(t)'' = -\omega_0^2 \sum (A_k \cos k\omega_0 t + B_k \sin k\omega_0 t); k \leq K \quad (3.1b)$$



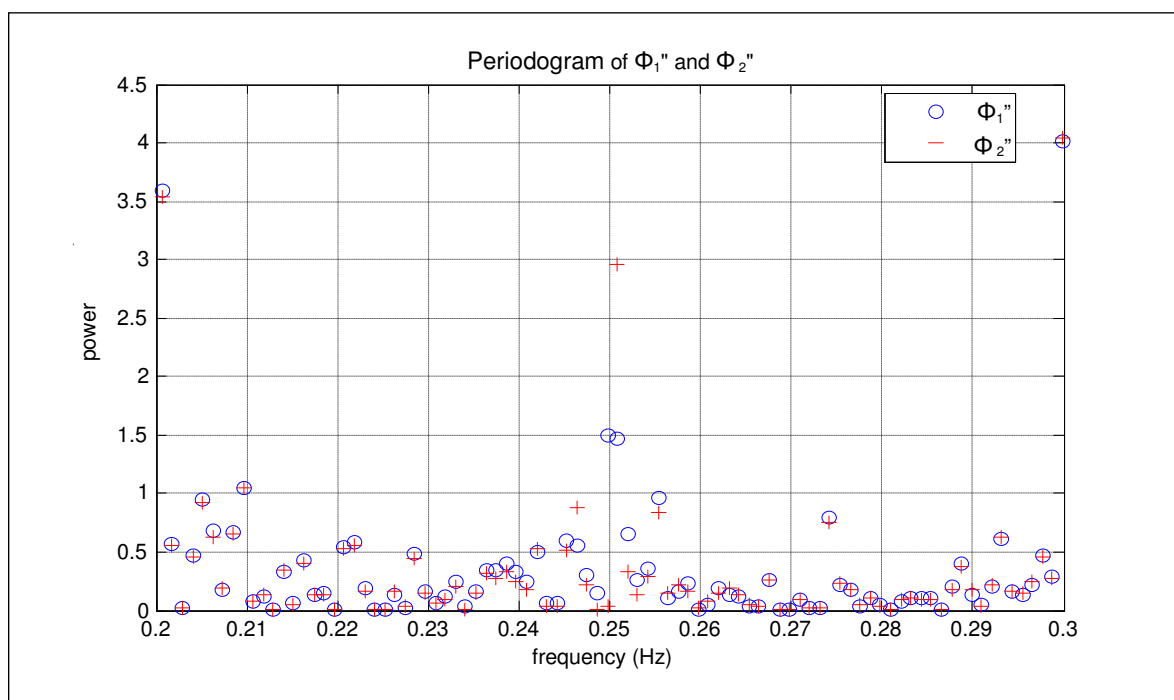
## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data

$$x_p(t) = \frac{A_0}{2} - \frac{x_p(t)''}{\omega_0^2} \quad (3.1c)$$

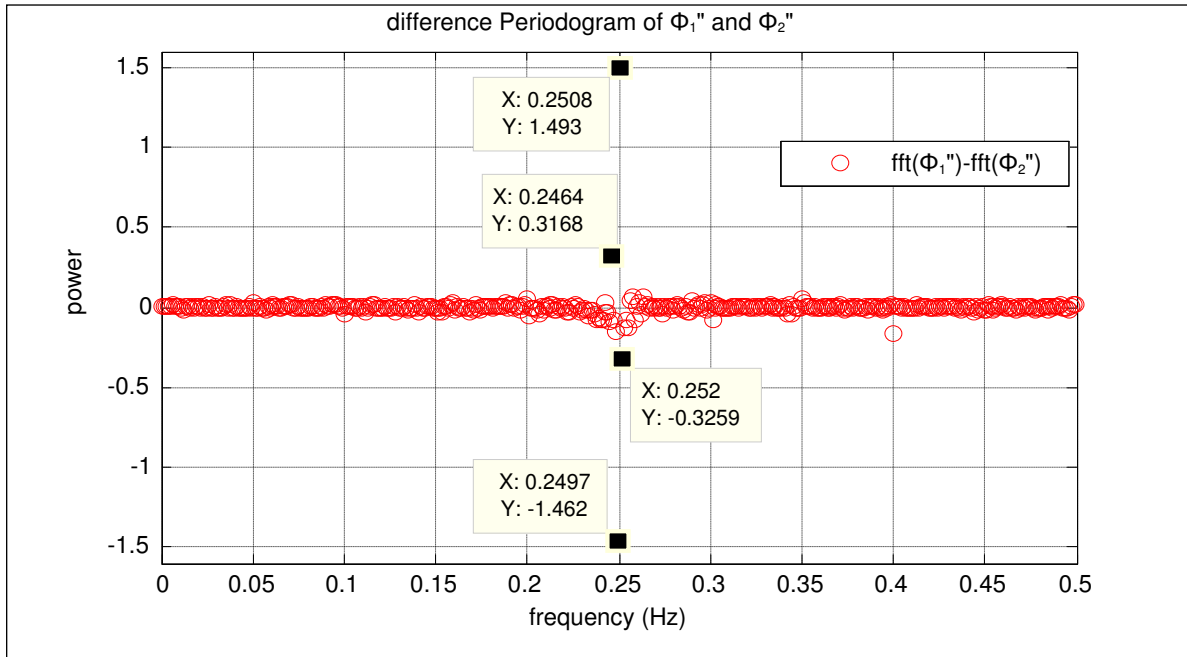
From discrete Fourier transform form (3.1a-c) we know, the derivative of  $\Phi_1$  or  $\Phi_2$  maintain their frequency information.



**Figure 4. 7 Second derivative of  $\Phi_1$  and  $\Phi_2$  (data from WUH2, doy 260 year 2007 PRN3)**



## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



**Figure 4.8 Periodogram of  $\Phi_1''$  and  $\Phi_2''$  and their difference (data from WUH2, doy 260 year 2007 PRN3)**

The above figure 4.8 is the periodogram of  $\Phi_1''$  and  $\Phi_2''$ . Close to the frequency 0.25 Hz there are two pairs of points are not matched. But it is still unknown, which carrier phase measurement causes the oscillation.

The three IGS stations, POTM in Potsdam, WUH2 in Wuhan and OBE2 in Oberpfafenhofen, have same receivers' and antennas' type: AOA BENCHMARK ACT and all of them are maintained by GFZ.

Other stations however with the same receivers Type, antennas Type and Responsible Agency have not the same problem. The reason is not jet clear. And after a discussion in GFZ, the 1 Hz GPS data from those three stations will be not used before the equipments are checked.

### 4.4 Single difference between satellites

From the above analysis we know the algorithm TurboEdit is not suitable to detect cycle slips in 1 Hz GPS data. And the single difference between satellites will be also tested. Since the orbiting GPS satellites transmit  $L_1$  at higher power than  $L_2$ ; as a result, the received  $L_1$  signal to noise ratio (SNR) will be better than that for  $L_2$ . Spatially, when the GPS satellites are in low elevation position, the noise at  $L_2$  is significantly bigger than the noise at  $L_1$ .

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data

When we combine  $L_1$  and  $L_2$ , the biggest of the noise in the combination comes from  $L_2$ . To take advantage of  $L_1$  signal the between-satellite single difference (one receiver and two satellites) of carrier phases will be more effective to detect cycle slips and outlier. Since the high rate data sampling can represent clearly the high variation in ionosphere delay, troposphere delay, and the most GPS receiver clocks can have high frequency variation, the between-satellite single difference eliminates the GPS receiver clock instability.

Flowing Figures 4.9-12 show us the data from PIE1 doy 260 year 2007

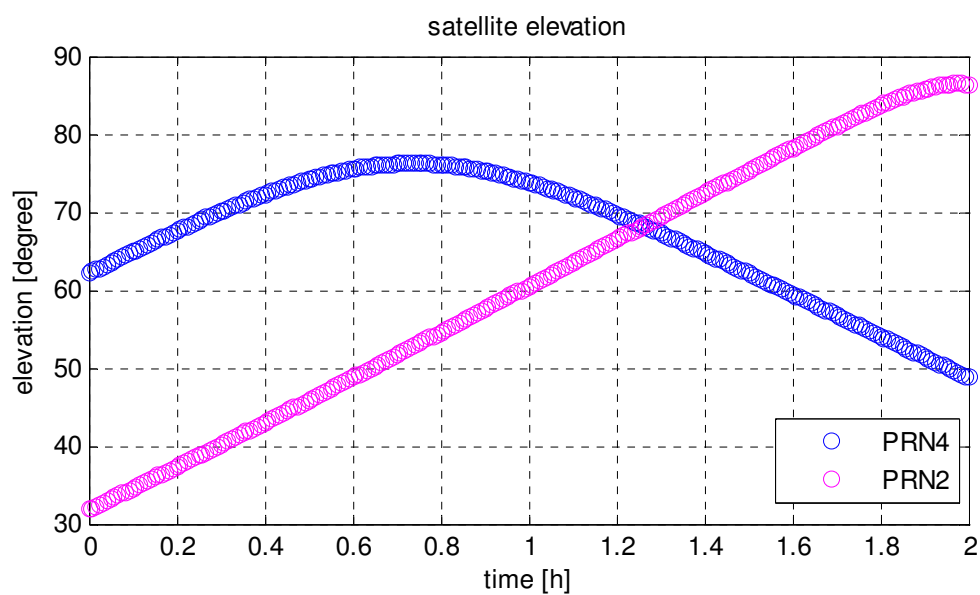


Figure 4.9 Two satellites elevation

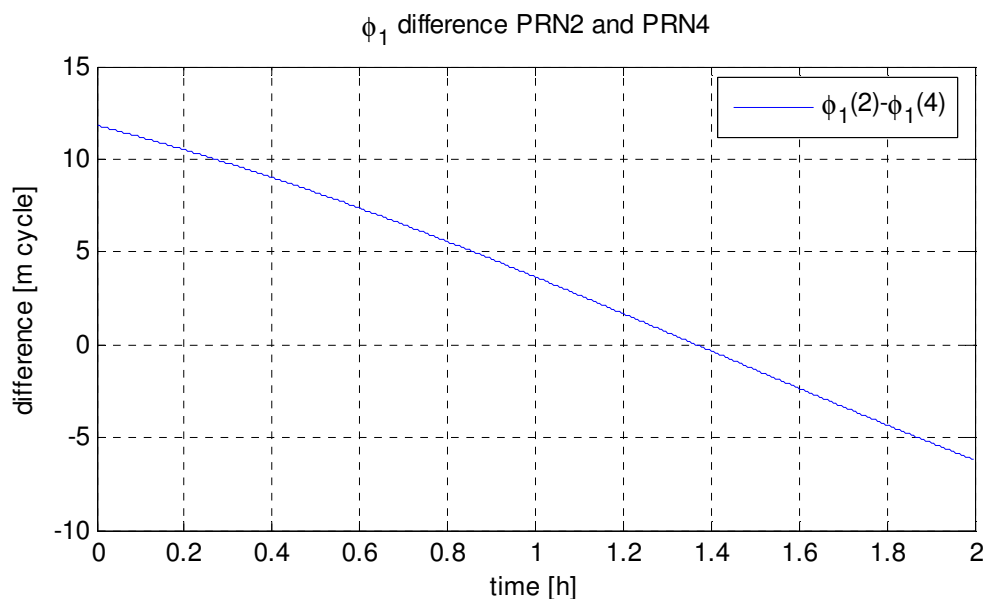
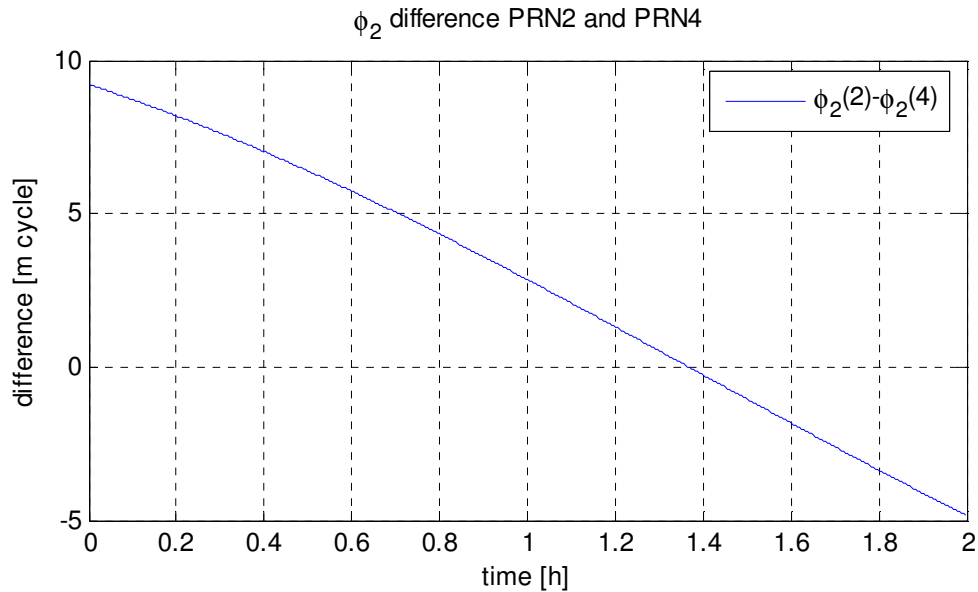


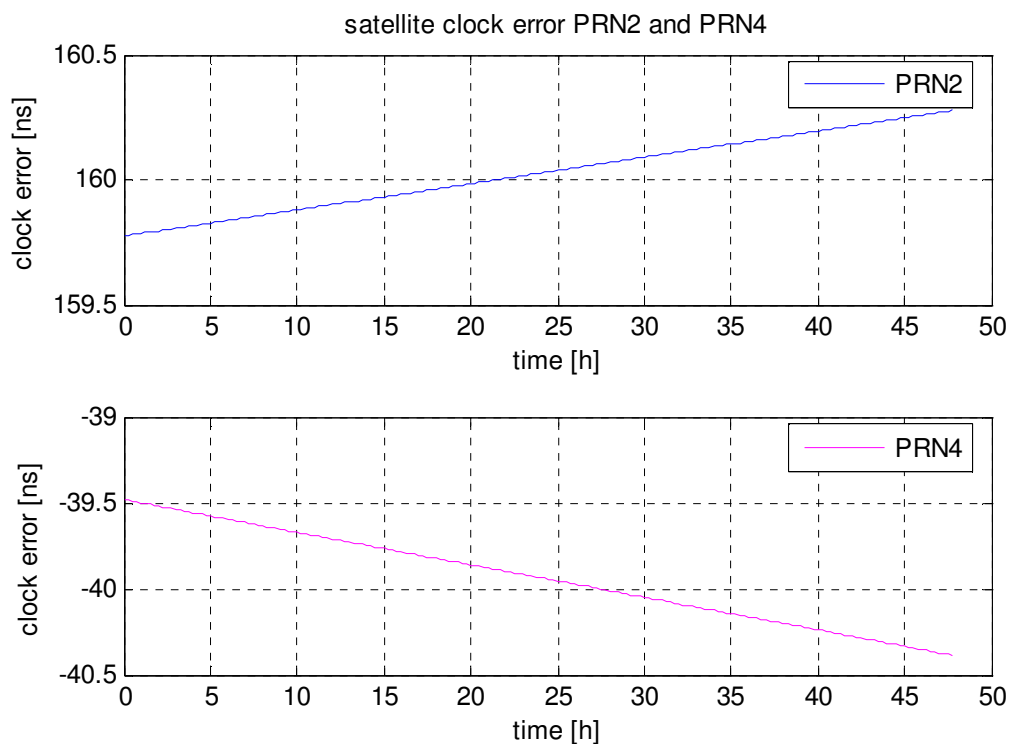
Figure 4.10  $\Phi_1$  single difference between PRN2 and PRN4

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data



**Figure 4.11  $\Phi_2$  single difference between PRN2 and PRN4**

The motion of a GPS satellite relative to other GPS satellite can be considered as uniform acceleration. It is the biggest contribution in the carrier phase single difference. Beside it the troposphere and ionospheric delay's difference between observations of two satellites is also a part of the carrier phase single difference. Those effects should be a reasonable smooth function.



**Figure 4.12 Satellites clock error of PRN2 and PRN4 (15 min data from SP3 file)**

## Chapter 4 Compare 1 Hz and 0.033 Hz GPS data

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The Figure 4.12 shows, that the satellite clock error is also a smooth function. The carrier phase single difference includes GPS satellites clock error difference. Since GPS satellites have an atomic clock, and it is much stable.

### 4.5 Conclusion

From the above analysis, we can get the follow conclusion: Because the precision of the pseudorange measurements in 1 Hz GPS data are not enough for the cycle slips detection, the TurboEdit algorithm can not be used for our purpose. After study the all the combination and difference, we found that the single differences between satellite are very smooth function, when there is no cycle slips occurred. With the single difference between satellites it will be very effective to detect cycle slips.

# 5. Detection with single differences

Base on the above analysis the single differences are implemented in the preprocessing procedure. In this chapter the computer program will be introduced. Since the goal of this work is to develop computer code for 1 Hz GPS data preprocessing, and like many other famous GPS software (Bernese, EPOS), the preprocessing module in GSEIS software package is also developed in Fortran 90 under Linux.

## 5.1 Input file and data streams

For the purpose of preprocessing the real time observation data stream  $L_1$  and  $L_2$  should be checked. To find out a reference satellite in possible short time, the satellites' elevations are sorted descending, and under this sequence the reference satellite is searched, which has no cycle slips and has possible little noise in its observations.

Since the rough estimated satellites' elevations are enough for our purpose. The approximate coordinates of GPS receivers can be gotten from the RINEX head file. To calculate elevation we only need the approximate coordinates of GPS satellites in the same coordinate system as GPS receivers. To get the satellites' approximate coordinates there are two ways, one is to compute them from the satellite information in the GPS broadcast ephemeris file; the other way is to interpolate the satellite coordinates directly from IGS products (predicted satellite orbit). Obviously the first way need more time than the second to calculate satellite coordinates, but it requests less communication with analysis centre to get IGS products.

$L_1$ and $L_2$	From GPS received raw data
GPS satellite position	From IGS products or estimated from GPS broadcast

**Table 5.1 Input data for preprocessing**

Because the GSEIS software is still in design phase, how to get the GPS satellite position in the preprocessing is still in question. In the current version the GPS satellite position is interpolated from IGS orbit product.

* 2007 12 20 9 15 0.00000000				
PG01	17400.865583	-19792.566411	3167.196289	173.713380
PG02	-7078.471709	14102.021908	21145.269210	159.777460
PG03	7836.401622	-23180.158410	-10134.872465	160.857400
PG04	-19523.553655	7593.772207	16451.426669	-39.487740
PG05	15978.084914	18235.146514	10745.534785	606.141380
PG06	17458.244004	2435.467297	20052.420923	183.624750
PG07	19040.245384	-1430.493386	18807.728881	-18.888550

## Chapter 5 Detection with single differences

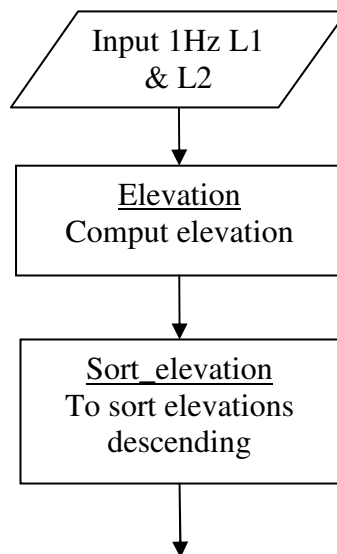
PG08	-24448.715209	812.570169	-11005.331229	-134.076320
PG09	15729.755304	14465.178097	-16500.558103	101.844530
PG10	56.360738	24534.473623	9621.838377	-179.577430
PG11	-18712.887891	-12473.534991	-14505.123545	26.682960
PG12	12305.971720	22444.411407	7318.287575	-351.926510
PG13	-17653.005267	-499.105950	19727.186544	229.063430
PG14	20962.971259	-14146.107325	-7853.852856	-311.801080
PG15	-32.713806	19235.590100	-18292.197058	-71.005570
PG16	3587.231153	-22248.166212	13994.796083	129.499990
PG17	-21080.229439	13411.783713	-8772.930604	46.190100
PG18	16404.283909	7788.268000	-19327.125977	-218.069100
PG19	888.377785	-18838.906904	-18554.159235	12.921960
PG20	-13851.075637	-20585.781150	9434.657676	124.163800
PG21	26242.758608	5636.866709	-1415.414275	74.983770
PG22	13576.129840	-8144.312873	-21161.182776	205.473080
PG23	-9068.963136	-12097.678352	21755.471983	319.462250
PG24	20872.208043	13045.647310	9770.653875	55.442720
PG25	-24870.449589	-7264.455352	5813.570416	492.783400
PG26	-3313.408035	19635.754750	-17738.816070	131.233000
PG27	-26330.982458	-4079.130036	-298.915788	155.361150
PG28	-14326.720425	4465.250189	-21554.254819	-12.325930
PG30	16639.154589	13166.449089	15705.174557	56.780570
PG31	17014.616976	-9267.217571	18409.872368	-2.091330

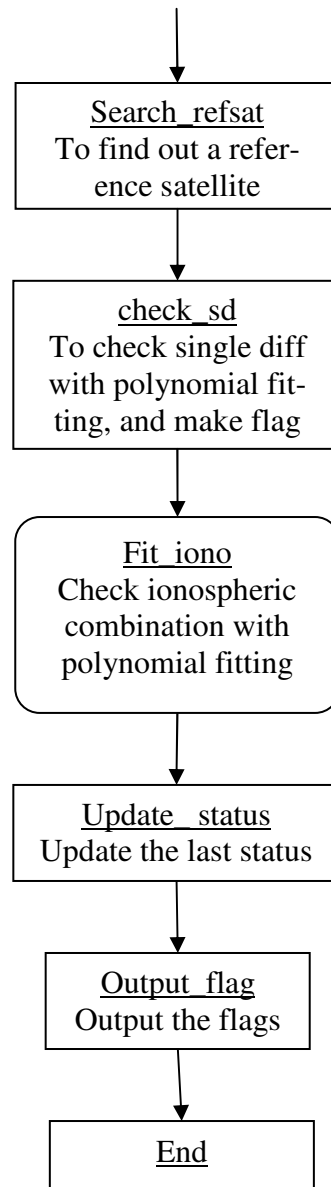
**Table 5.2 IGS orbit product in SP3 format**

The first line is the GPS time of the prediction. In the following lines there is the satellites' number, components of position in meter and clock bias in nanosecond.

### 5.2 Structure of the software

In every second the program gets GPS observations from the NTRIP data streaming, after processing the observations, which are tested as cycle slip or outlier, are flagged, and the flags will be outputted for the parameter estimation.





**Figure 5.1** Main program diagram of `chn_filter`

During the preprocessing the last 10 seconds carrier phase observations from  $L_1$  and  $L_2$  are always updated, which are prepared for the next polynomial fitting. Every second the program estimates the newest coefficients for all satellites' observation, it can deal with high ionosphere activity very effectively, and it is flexible to change reference satellites. Compared to radio signal the GPS signal is very weak, receiver can loose lock GPS satellite signal for a short time, from several seconds to several minutes, after that the ambiguity is still possibly unchanged. For this case the last estimated polynomial coefficients are also saved with reference satellite number and reference time together for every satellite. When the signal loss is longer than 10 minutes, the saved polynomial coefficients of the satellite are set invalid.

In the program there are two important steps, initialization and single difference checking.



## Chapter 5 Detection with single differences

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- a) In initialization phase, if the reference satellite can not be found, the program must wait for enough observations without cycle slips or outlier.
- b) In cycle slip detection, the last 10 single differences between satellites are fitted with a quadrate polynomial, which is estimated with the least squares method.

### 5.3 Initialization phase

To make single difference between satellites we need a reference satellite, which has low noise in observations  $L_1$  and  $L_2$  at the last 10 epochs. This means that after the program start or after a long time without GPS data, it needs at least 10 seconds to initialize.

To save time in this phase the program searches for reference satellite in descending order of the satellites elevations' angles. Since the GPS observations from higher elevation angle have less noise than those from lower elevation. And we do not have to check  $L_1$  and  $L_2$  separately but only the ionospheric combination from  $L_1$  and  $L_2$ .

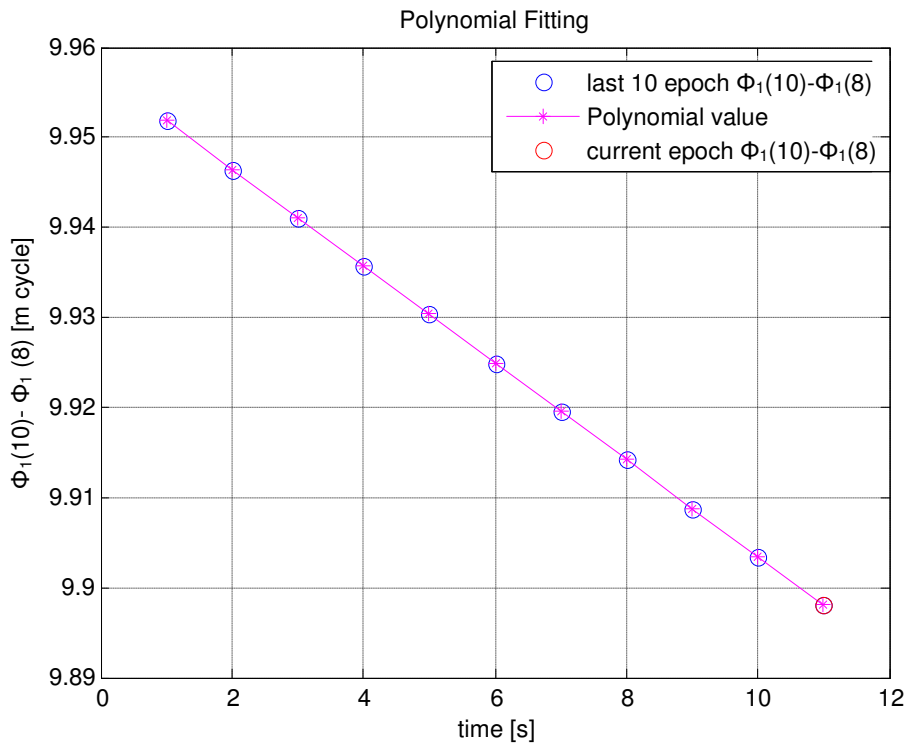
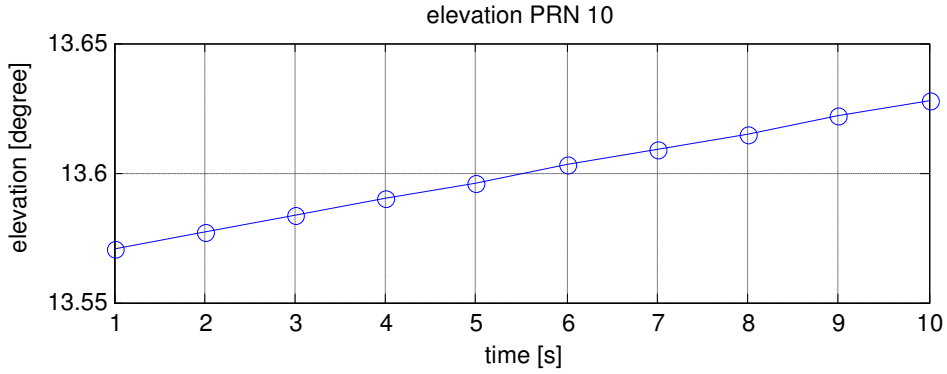
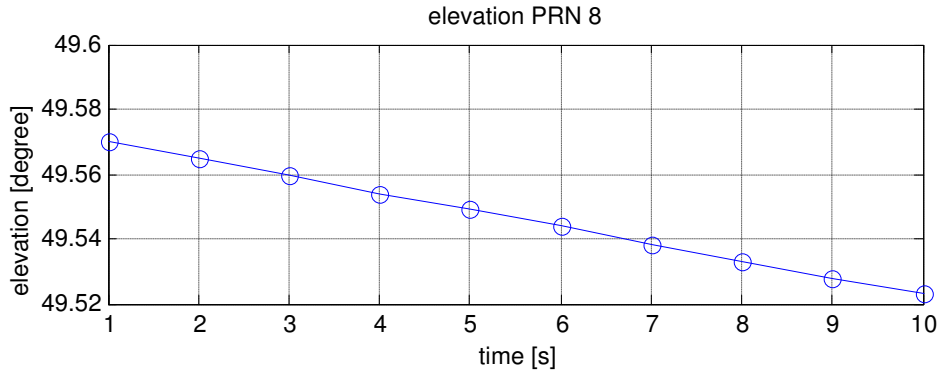
When a satellite has 10 available observations in  $L_1$  and  $L_2$  in the last 10 seconds, and after polynomial fitting of the ionospheric combinations the standard deviation RMS is also smaller than 0.3 cycles. Then we extrapolate the estimated polynomial to current epoch. If the difference between the ionospheric combination in the received current epoch and the extrapolated value is smaller than 0.6 cycles, then this satellite is suitable as reference satellite for the current epoch. Otherwise the next satellite will be tested. When we can not find a reference satellite after checking all of satellites, the program has to wait until enough observations without cycle slip or outlier are received.

When a GPS satellite arises, the observation in first 5 epochs will not be flagged. Then the program will try to fit them with 2 degree polynomial. If the polynomial fitting is successful, then the initial phase of the GPS satellite data is finished, otherwise the program must wait for more observations from this new GPS satellite.

### 5.4 Polynomial Fitting

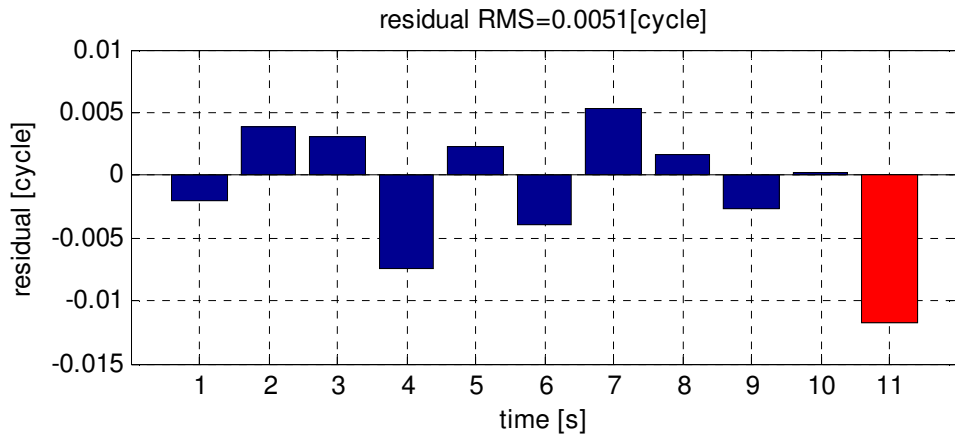
After the successful initialization the  $L_1$  and  $L_2$  single difference between reference satellite and other satellites in the last 10 seconds are calculated. They will be checked for cycle slips separately.

# Chapter 5 Detection with single differences



## Chapter 5 Detection with single differences

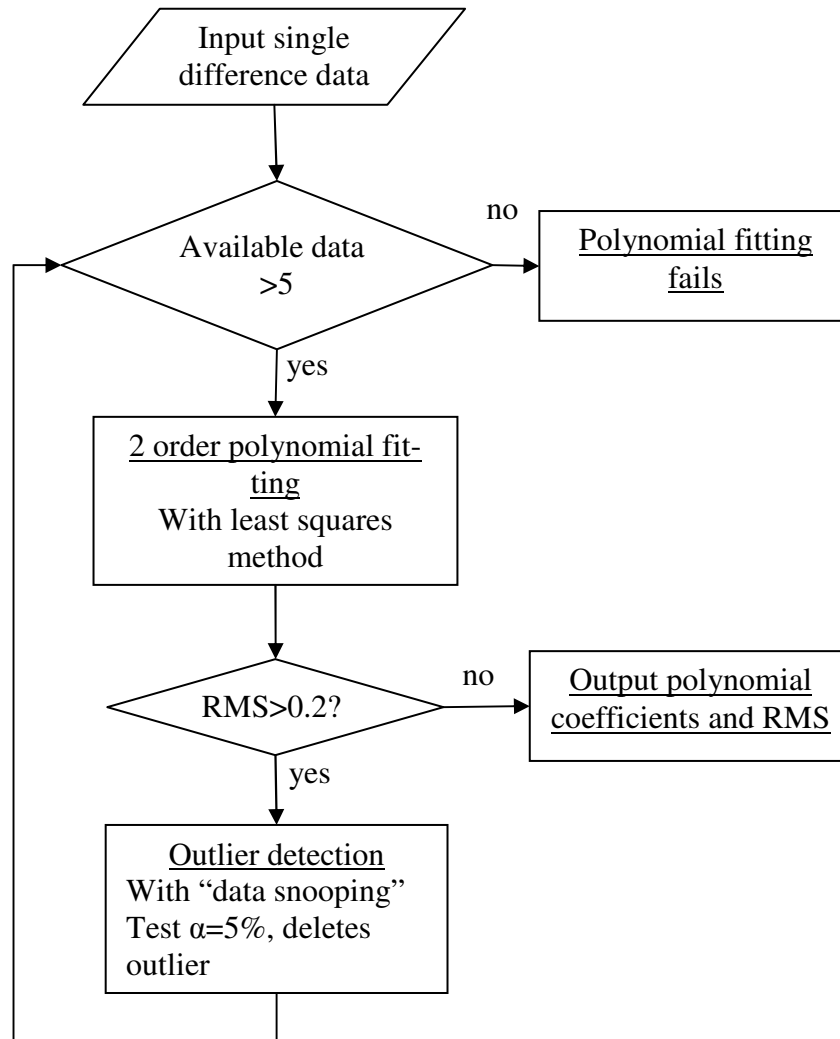
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**Figure 5.2** Elevation angles of two satellites (top)  $L_1$  single difference polynomial fitting (middle) and residual (bottom)

In Figure 5.2  $L_1$  differences of the last 10 epochs between the reference satellite 8 and satellite 10 is fitted with a 2-order-polynomial. The RMS (Root Mean Square) shows the quality of the last 10 epochs  $L_1$  and the estimated polynomial. From the residual of the current epoch we can find out a cycle slip.

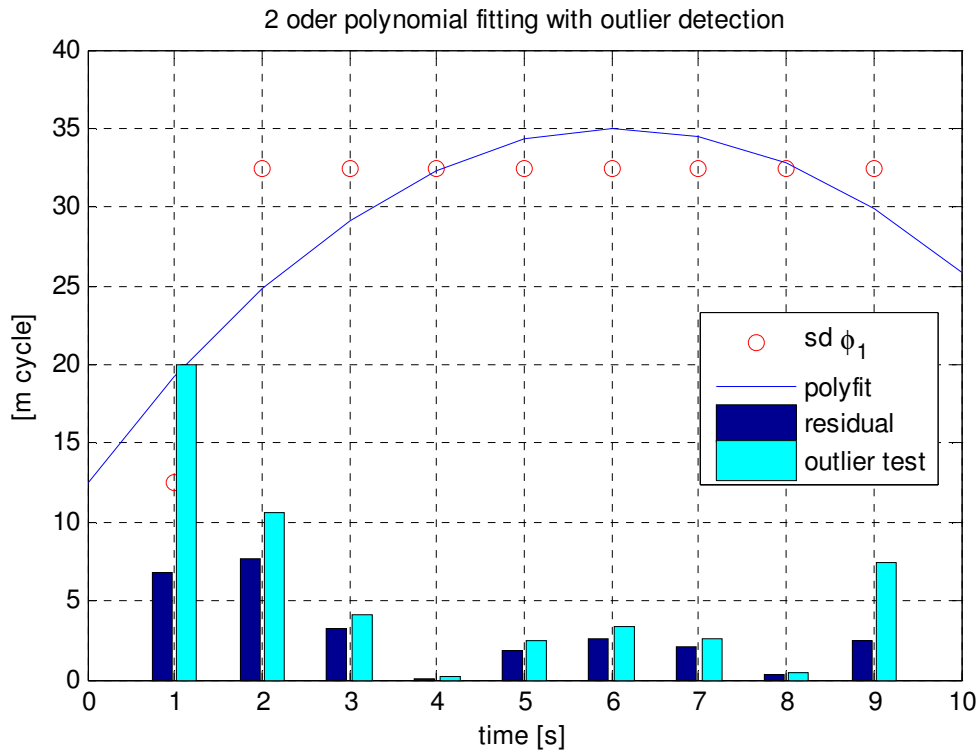
When there are some outliers or observations with big noise in the last 10 epochs, they will not be used in the polynomial fitting. For this reason the flags of the last 10 epochs' observations are also saved in a matrix, 1 for observations with big noise or outlier, 0 for good observations. But at the initialization phase we give 1 to the flags of the last 10 epochs without checking. This means, the polynomial fitting should be able to estimate coefficients robustly. In this program when the RMS is too big after 2 order polynomial fitting, and there are still more than 5 good observations in the last 10 epochs, then a one order polynomial is estimated and the data with the largest residual is detected as outlier, with rest data the next try of 2 order polynomial fitting begins.



**Figure 5.3 Outlier detection in polynomial fitting**

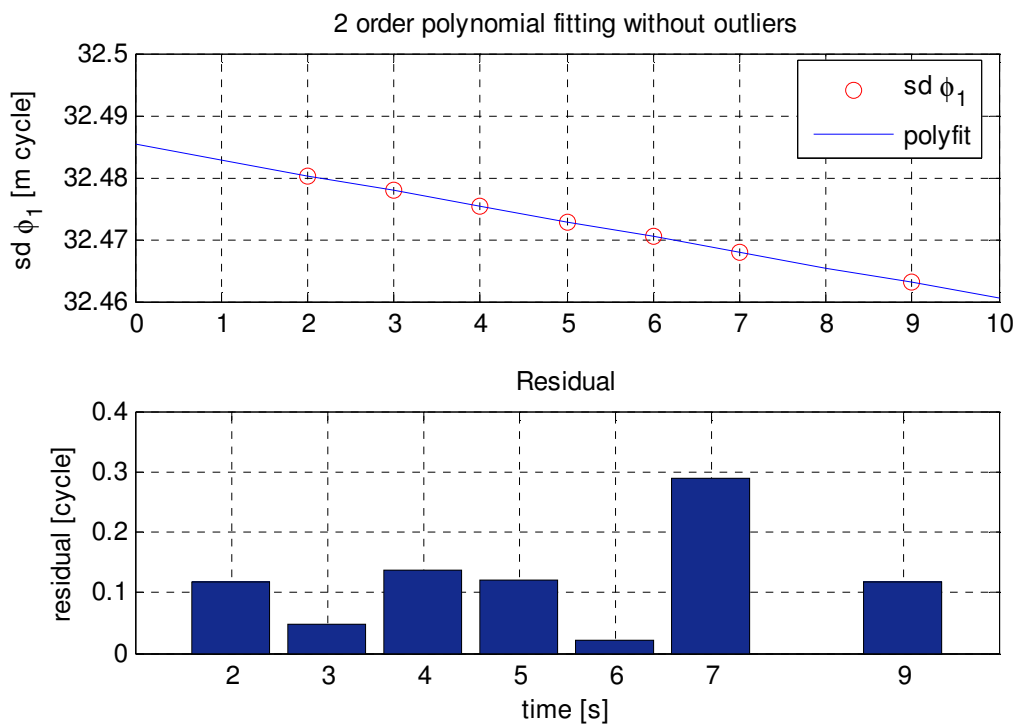
Figure 5.3 shows polynomial fitting. If the polynomial fitting fails, the program will wait for next epoch data, until polynomial fitting is successful.

## Chapter 5 Detection with single differences



**Figure 5.4 outlier detection with 2 order polynomial fitting (sd: single difference)**

In this example there are two outliers in time 1 and 8. They will be detected by outlier test POPE.



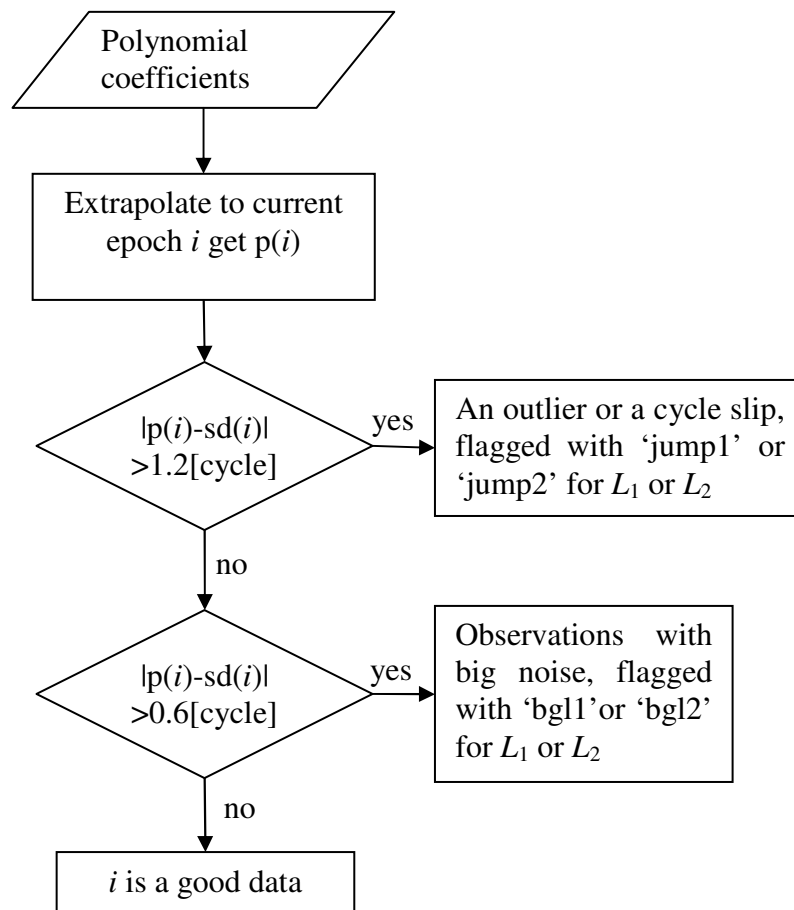
**Figure 5.5 2 order polynomial fitting without outliers**

## Chapter 5 Detection with single differences

Because the last 10 epochs with outliers happen often in initial phase, the outliers must be cleared before polynomial fitting.

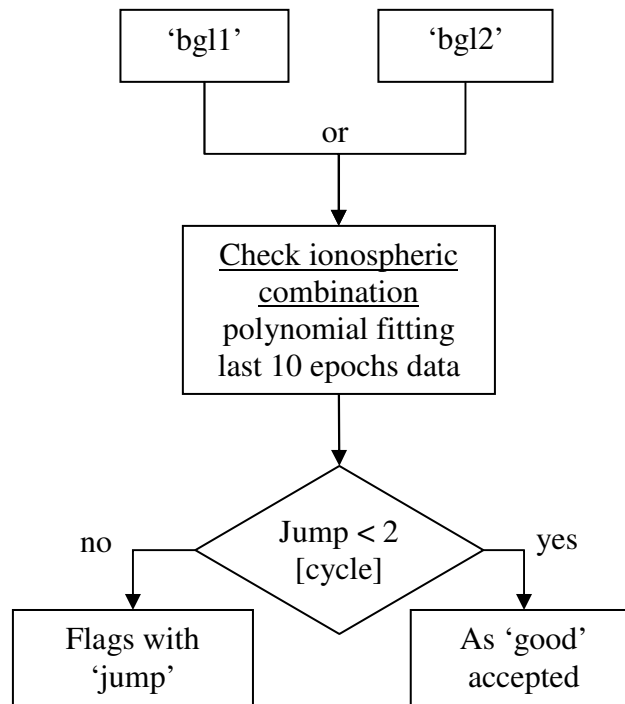
### 5.5 Single difference checking

When the single difference between the reference satellite and an other satellite in the last 10 epochs is fitted with a 2 order polynomial successfully, or the last successful polynomial fitting is not 10 minuses earlier than current epoch, then the polynomial is extrapolated to current epoch to get a model value; the single difference of current epoch is required to lie within 0.6 cycle of the extrapolated value.



**Figure 5.6** check current epoch

Carrier phase  $L_1$  and  $L_2$  should be checked separately. After the check there are 3 possible results: 'jump', 'big noise' and 'good data'. For the result 'big noise', there is possible one cycle slip occupied. One cycle slip in carrier phase  $L_1$  or  $L_2$  can causes bigger than 3 cycles jump in their ionospheric combinations. Thus we can check the ionospheric combinations with polynomial fitting, when a 'big noise' is flagged in carrier phase  $L_1$  or  $L_2$ .



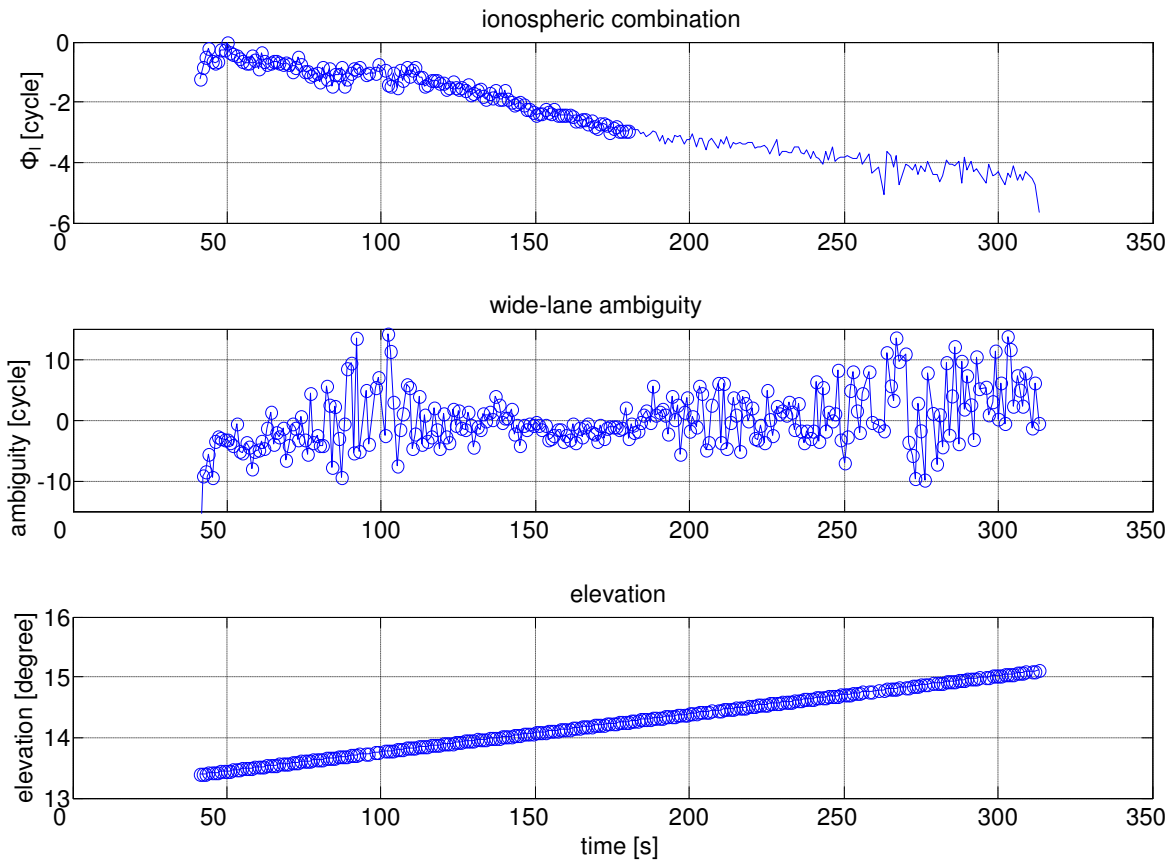
**Figure 5.7 Check big noises with ionospheric combination**

Because of its smaller variation the ionospheric combination is very effective to check the carrier phase with big noise.

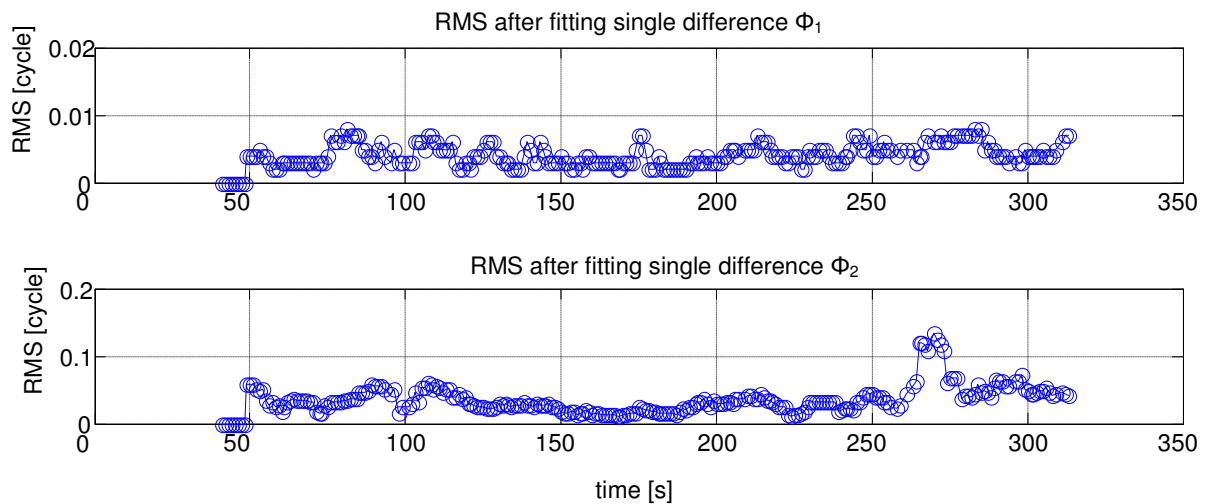
### 5.6 Results

To test this program the 6 days data from about 110 IGS GPS stations are downloaded and checked. Results are in the following figures.

## Chapter 5 Detection with single differences



**Figure 5.8 Variations in combination of 1 Hz GPS data**



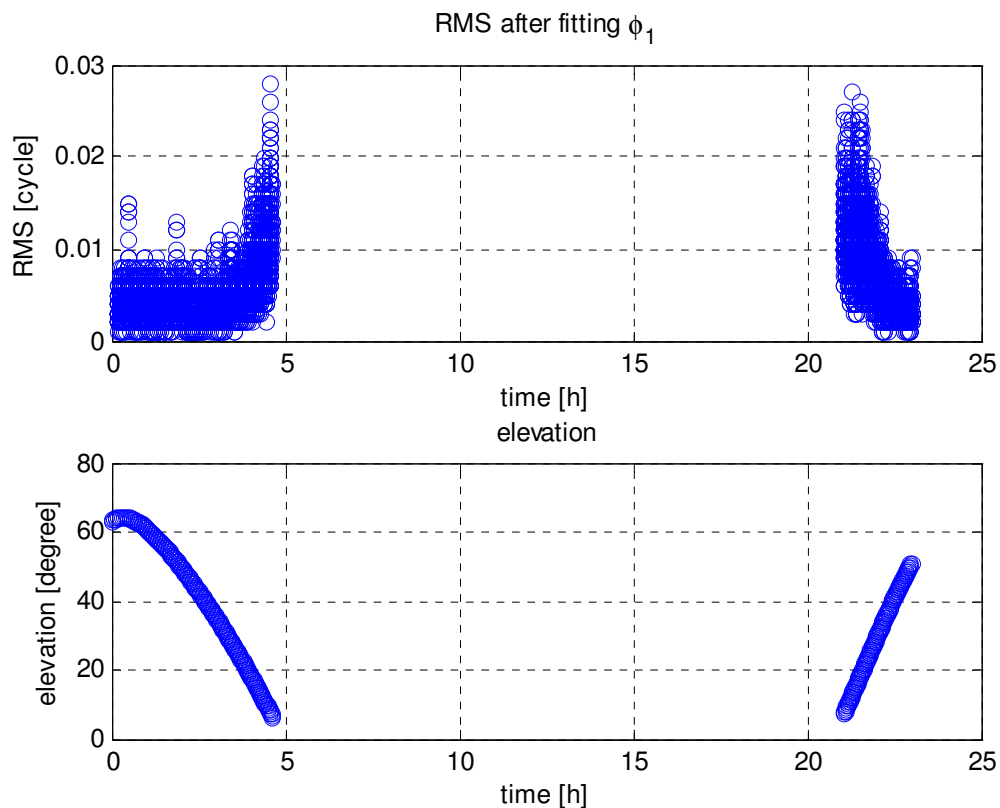
**Figure 5.9 RMS after polynomial fitting of 1 Hz GPS data**

In both figures 5.8-9 we used the data from ULAB doy 260 year 2007 PRN11. The same 1 Hz GPS data is checked with difference method. In the Figure 5.8 the ionospheric combinations have a large than 1 cycle jump at time = 264; it can be possible that both of  $L_1$  and  $L_2$  have one cycle slip. And from the wide-lane combinations we can not get any help. In this case the two combinations are not suitable to detect cycle slip. As discussed in the previous chapter, which the RMS of polynomial fitting identifies, the carrier phases quality in the last 10 ep-



## Chapter 5 Detection with single differences

ochs. From the figure 5.9 the RMS in  $L_2$  is about 10-20 times bigger than the RMS in  $L_1$ , and the bigger variation in  $L_2$  is the cause of the variation in ionospheric combination.



**Figure 5.10 RMS and elevation angle**

In figure 5.10 the 1 Hz GPS data of the station PIE1 in doy 260 year 2006 PRN1 are used. The experiments show a high negative correlation between satellite elevation angle and RMS of polynomial fitting. In these experiments the cutoff elevation angle is 7 degrees. Even at low elevation angle the RMS are still small enough to check out the cycle slips with only one cycle.

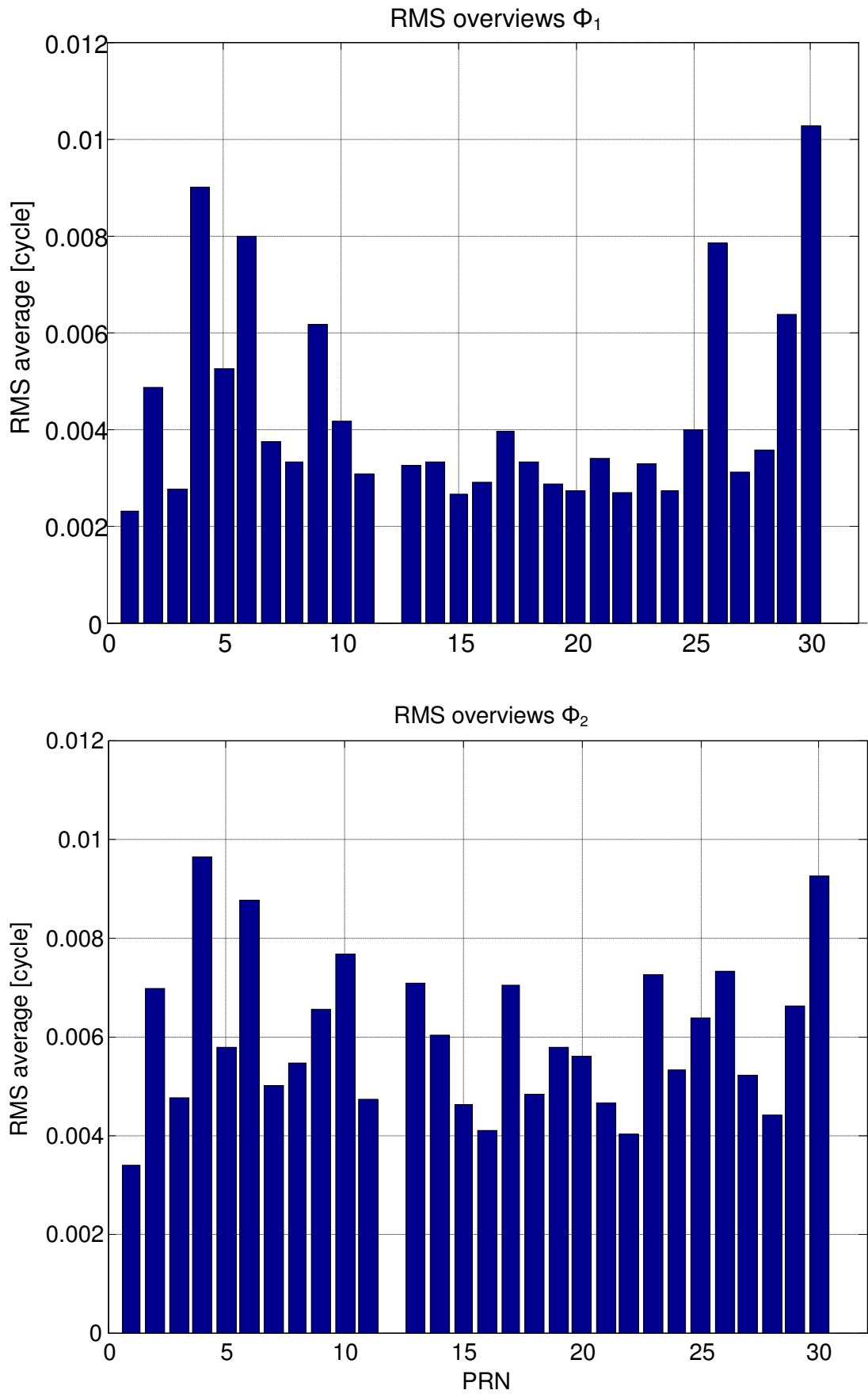


Figure 5.11 RMS after polynomial fitting for all satellites in one day

## Chapter 5 Detection with single differences

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doy year	Elevation $\geq 7^\circ$	good $\Phi_1$	%	good $\Phi_2$	%
100 2006	733461	731375	99.7	731375	99.7
200 2006	706329	703989	99.7	703988	99.7
260 2007	683967	682538	99.8	682538	99.8
320 2007	758406	756574	99.8	756575	99.8

**Table 5.3 Rate of good data after checking (station: PIE1)**

Table 5.1 is the checking result of the 1 Hz GPS data from IGS station PIE1. The numbers in Table 5.1 the 2nd column are the sums of the observations with higher than 7 degrees elevation in the test day for all satellites. And in the 3rd and 4th column there are the sums of carrier phase observation without cycle slip or outlier after Channel\_Filter check. The most lost data are caused by initialization phase.

After checking 6 days of about 110 IGS stations with the program Channel\_Filter, the average rate of good data  $\Phi_1$  is 99.59 %, and for  $\Phi_2$  it is 99.55 %.

# 6. Analysis of single frequency receivers

Since single frequency code/carrier phase GPS receivers are less expensive than dual frequency receivers, and many single frequency solutions exhibit equivalent accuracy as those obtained from dual-frequency observations, single frequency receivers become more popular nowadays. GFZ plans to build a GPS network with 1500 single frequency receivers in Germany for meteorology research. Here the receiver clock bias and carrier phase measurement of three single frequency GPS receiver (GARMIN, THALES and NOVATEL) are tested and compared. During which the Channel\_filter is used.

## 6.1 Determination of reference satellite

For single-frequency carrier phase GPS receivers it is impossible to determine a reference satellite with the ionospheric combination. Since GPS observations from higher elevation angle have less noise than those from lower elevation. Here we set the satellite with the highest elevation as the reference satellite.

## 6.2 GARMIN (GPS 17-HVS)

The first tested GARMIN (GPS 17-HVS) is a low cost single frequency GPS receiver, which outputs raw pseudorange and carrier phase data. It can track and use up to twelve satellites with a default data rate of 1 Hz in the Garmin proprietary format. After conversion to standard RINEX format, we get:

08	01	21	01	13	12.2044286	0	7G08G28G18G09G26G17G15
-514001149.557	3566712596.295	42.000					
-516342019.704	3839321302.908	51.000					
-515349264.667	4093958455.476	36.000					
-516401070.332	4145103917.072	47.000					
-517824036.326	3849405565.825	52.000					
-515671088.193	4144350802.482	50.000					
-517834638.842	3862259609.821	53.000					
$C_1$	$\Phi_1$	$S_1$					

**Table 6.1 One epoch 1Hz GPS data from GARMIN (in RINEX form)**

Usually the distance between a satellite and a site is more than 20,000 km, in the above received observations the  $C_1$  is negative. This means that the clock of GARMIN GPS 17-HVS

## Chapter 6 Analysis of single frequency receivers

has big offset. Before we begin to preprocess the 1Hz GARMIN GPS data, the clock offset must be estimated. Here we attempt to calibrate the receiver clock using GPS broadcast ephemeris.

$$C_1^k = \rho^k + dt^k + dt + I^k + T^k \quad (6.1a)$$

$$dt + I^k + T^k = C_1^k - \rho^k - dt^k \quad (6.1b)$$

$I^k$  : ionospheric delay

$T^k$  : troposphere delay

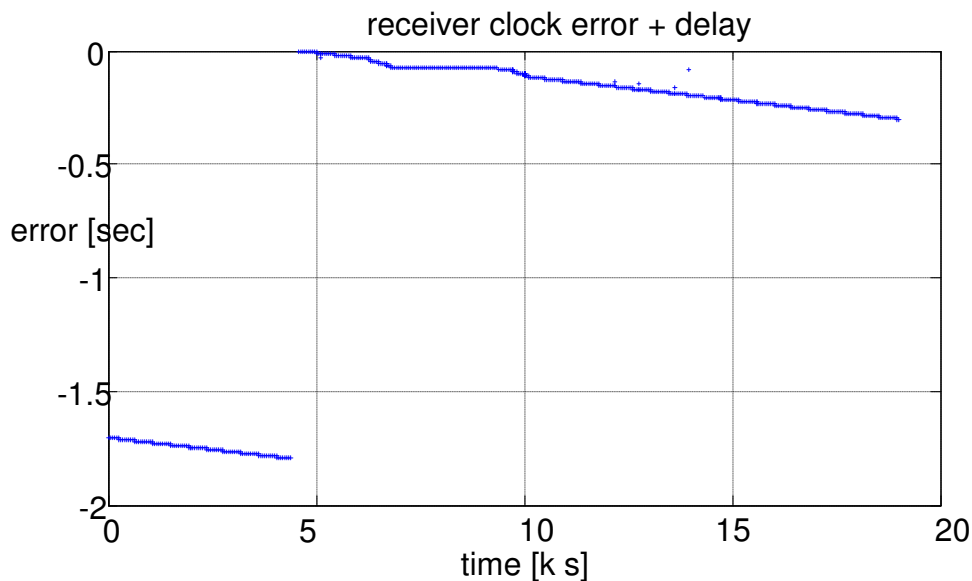
$\rho^k$  : distant between satellite and site, calculated from GPS broadcast and the known site coordination

$dt^k$  : satellite clock error, calculated form GPS broadcast

$dt$  : receiver clock error

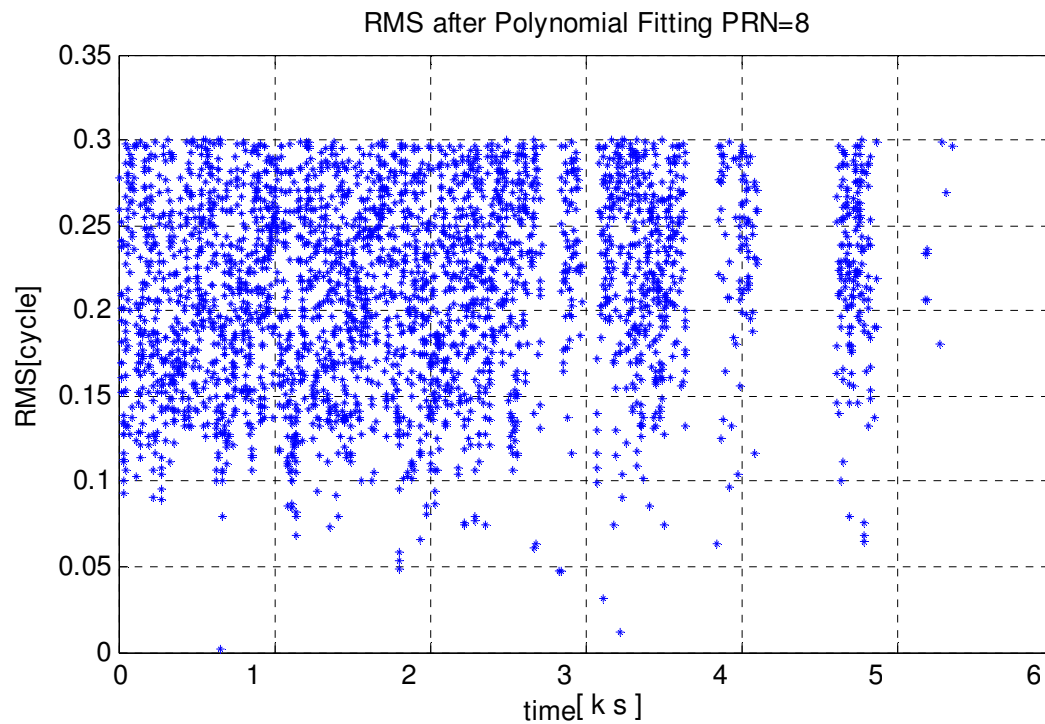
$k$  : satellite index

From equation (6.1b) we can get the sum of receiver clock error, ionospheric and troposphere delay.



**Figure 6.1 Receiver clock errors, ionospheric and troposphere delay**

At time=4.6 [m s] the GARMIN receiver clock was synchronized with an atomic clock. The GARMIN receiver clock has not only offset, but also drift. After correcting the receiver clock error and regardless of the delay the Channel\_filter is run to check the RMS after polynomial fitting.



**Figure 6.2 RMS for GARMIN 1 Hz data checking**

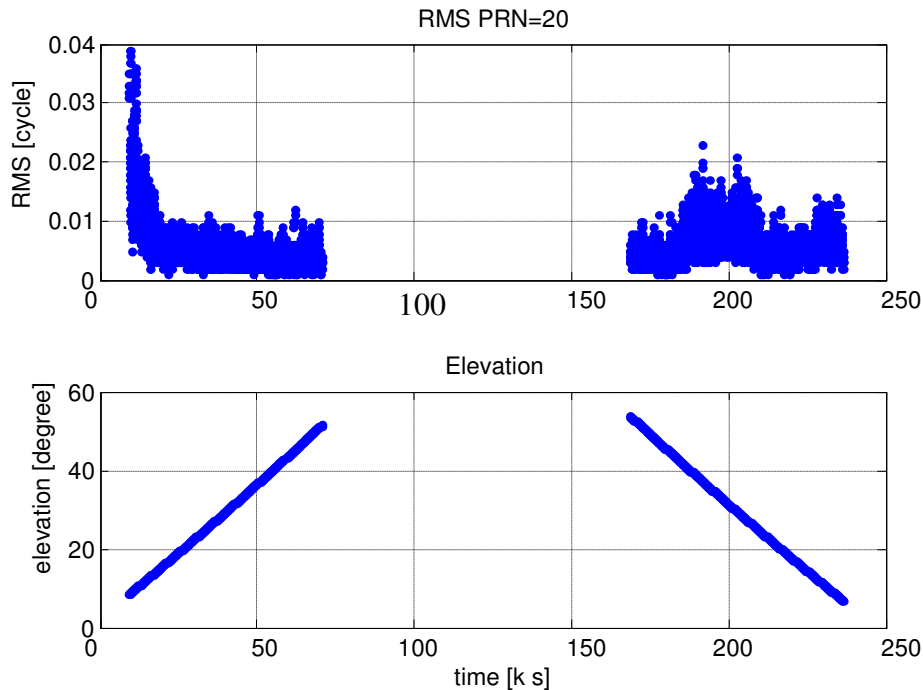
The RMS of GARMIN carrier phase polynomial fitting is disorderly, and it is impossible to detect a cycle slip. GARMIN (GPS 17-HVS) can not be used for GPS carrier phase observation.

### 6.3 THALES (AC-12)

The THALES AC12 receiver provides raw data output with carrier phase measurements. With THALES(AC-12) 3 days of 1 Hz GPS data have been generated. After checking raw data, there is no conspicuous offset and drift of the receiver clock.

With the program Channel\_filter the RMS of polynomial fitting will be checked.

## Chapter 6 Analysis of single frequency receivers



**Figure 6.3 RMS for THALES 1 Hz GPS data checking**

In the gap of Figure 6.3 the PRN20 was taken as reference satellite.

doy year	Elevation $\geq 7^\circ$	good L1	%
034 2008	796653	794080	99.7
035 2008	798109	797403	99.9
036 2008	796704	795777	99.9

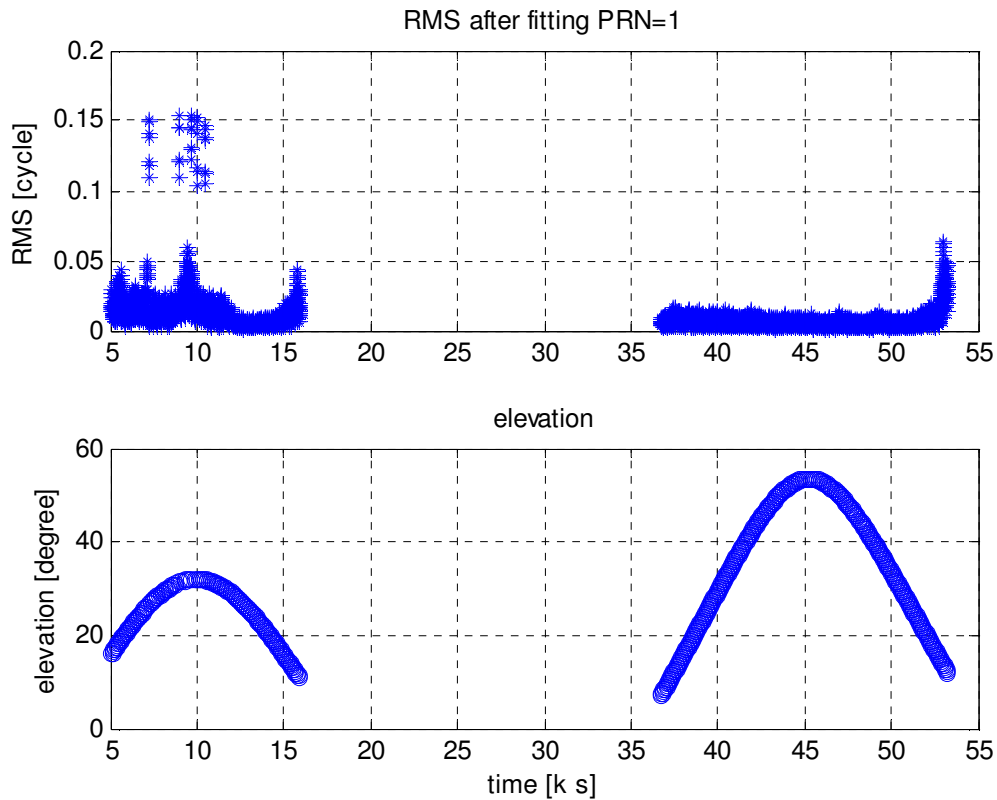
**Table 6.2 Rate of good data after checking (THALES)**

This study showed that the quality of the GPS carrier phase measurements of the single-frequency THALES (AC 12) is equivalent to the dual-frequency GPS receiver.

### 6.4 Novatel (smart antenna<sup>TM</sup>)

The SMART ANTENNA<sup>TM</sup> integrates a GPS receiver and an antenna; it features 12-channel code and carrier phase tracking for improved positioning accuracy and reliability. It provides position, velocity, and time (PVT) output at rates up to 5 Hertz or raw carrier phase measurement data at rates up to 10 Hertz. [novatel]

## Chapter 6 Analysis of single frequency receivers



**Figure 6.4** RMS for Novatel 1 Hz GPS data checking

The RMS of Novatel has some jumps at the satellite at big elevation, but it is correlated with elevation.

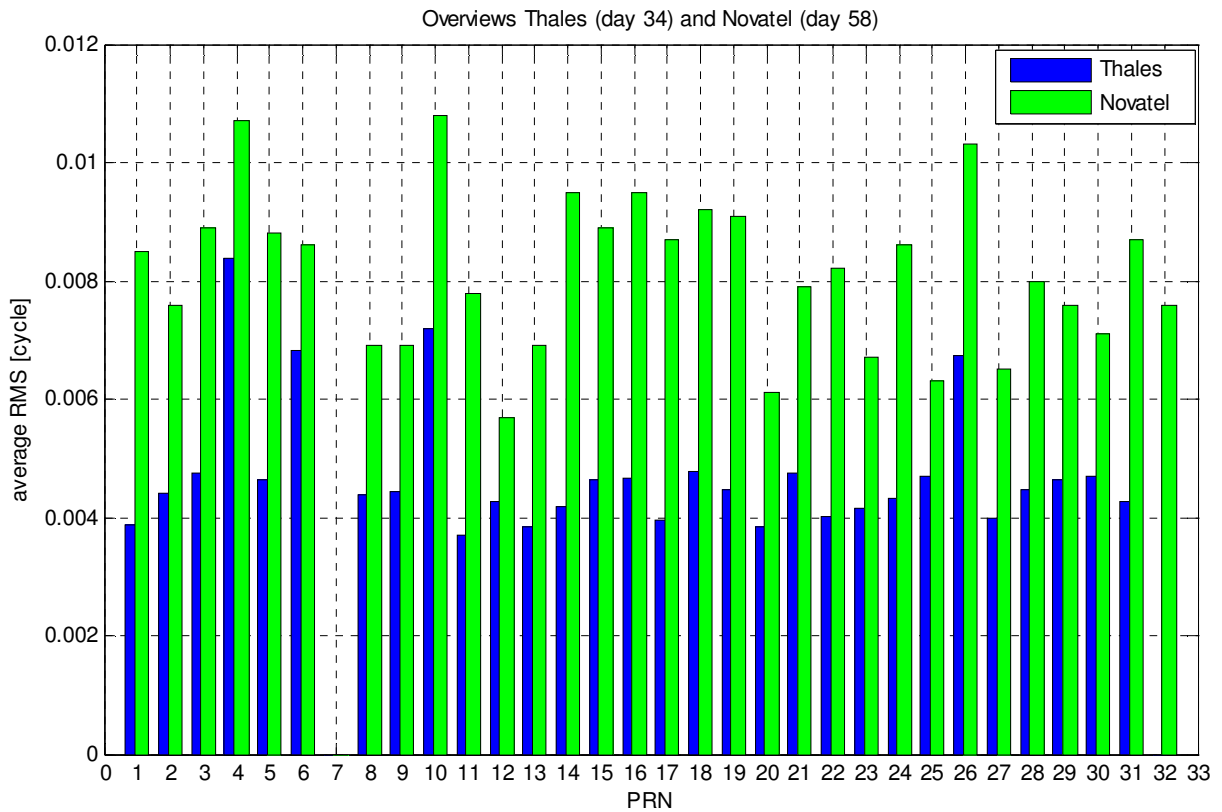
doy year	Elevation $\geq 7^\circ$	good L1	%
054 2008	112629	112310	99.7
058 2008	796704	795777	99.65

**Table 6.3** Rate of good data after checking (Novatel)

At the doy 054 there were only 4.25 hours observation data.



## Chapter 6 Analysis of single frequency receivers



**Figure 6.5 RMS after polynomial fitting for all satellites (THALES & NOVATEL)**

Comparing THALES with NOVATEL the noise in the THALES received GPS data is smaller than in NOVATEL received GPS data.

# 7. Summary and conclusion

- a) In this thesis a new 1 Hz GPS data preprocessing method for real time applications is proposed and implemented in Fortran 90. The preprocessing module, Channel\_filter will be used in GSEIS project to detect cycle slips and outliers in 1Hz GPS data stream. Because the method based on PPP real-time positioning is used in GSEIS project, the preprocessing should be able to be used for single GPS station.
- b) Currently the algorithm TurboEdit is used publicly in PPP preprocessing for single GPS station. After study the TurboEdit algorithm and EPOS software the new method Channel\_filter and its FORTRAN software was developed to detect cycle slip and outlier in real time 1 Hz GPS data. The Channel\_filter algorithm is documented by many figures in Chapter 5.

For the examination six days 1 Hz dual frequency GPS data was downloaded from IGS data center and simulated as real-time data stream. In the TurboEdit algorithm there are two combinations, ionospheric and wide-lane combination, which were used for cycle slip or outlier detection. During the examination of the GPS data it was found that the 1 Hz GPS data pseudorange measurements are less precise than the 0.033 Hz GPS data. And the wide-lane ambiguity has very large scatter, and can not be used to detect cycle slips or outliers. Because the cycle slips can occur in  $L_1$  and  $L_2$  concurrent in different value, at least two combinations are need to detect cycle slips in  $L_1$  and  $L_2$  other effective detection will be found out to take place of wide-lane combination.

- c) Since the GPS satellites transmit  $L_1$  at higher power than  $L_2$ ; as a result, the received  $L_1$  signal has less noise than  $L_2$ . To take advantage of the  $L_1$  signal the between-satellite single difference (one receiver and two satellites) of carrier phases are a better choice. With this idea the Channel\_filter algorithm is developed. In the Channel\_filter program there are three steps to detect cycle slip or outlier:
- Find a reference satellite with enough good observations in the last 10 epochs.
  - Compute the last 10 epochs  $L_1$  and  $L_2$  single differences between the reference satellite and the other satellite, and then fit a second-order polynomial to the differences.
  - Extrapolate the second-order polynomial to the current epoch, and check the residuals. If a residual is bigger than a certain limit, possibly there is cycle slip or outlier in the current epoch.

## Chapter 7 Summary and conclusions

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- d) To make a successful detection the RMS of the second-order polynomial fitting should be small enough to detect one cycle slip. In a test with the 6 days 1 Hz GPS data of 110 stations the calculated RMS are less than 0.1 [cycle] for  $\Phi_1$  and 0.2 [cycle] for  $\Phi_2$ . It is easy to check out one cycle slip or outlier with so small RMS.
- e) Since single-frequency code/carrier phase GPS receivers are applied in many real-time projects, in this thesis the 1 Hz GPS data of three single frequency GPS receivers GARMIN, THALES and NOVATEL are checked with Channel\_filter.
- The GARMIN GPS receiver clock has a drift about 1.7 second/day, and the carrier phase measurements are irregular, and can not be applied for positioning.
  - The THALES and NOVATEL GPS receiver exhibit equivalent carrier phase observations as due frequency GPS receiver.

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